

# STATISTICAL HADRONIZATION MODEL FOR HEAVY-ION COLLISIONS AT SIS18 ENERGIES

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PRC 00, 004900 (2023) (to appear)



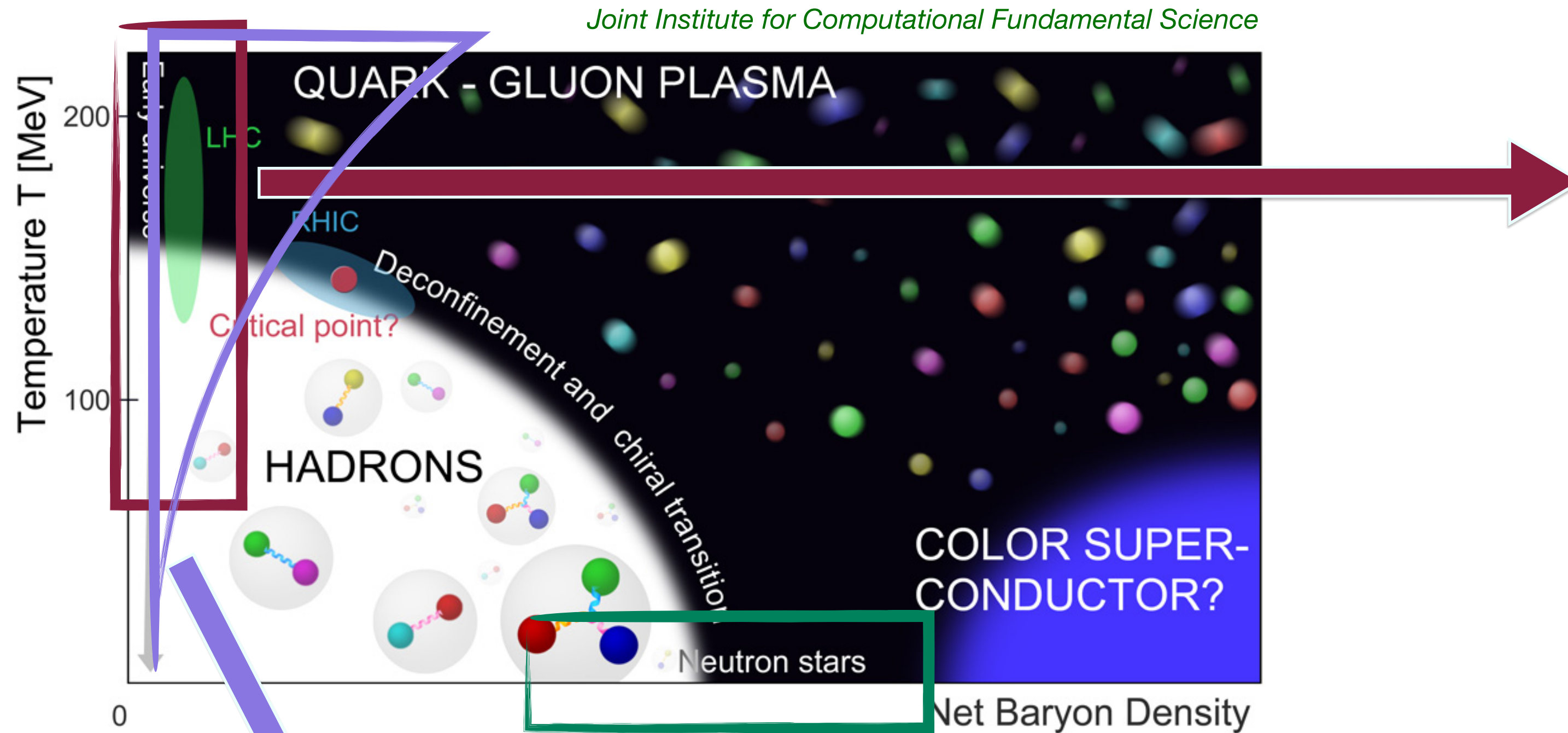
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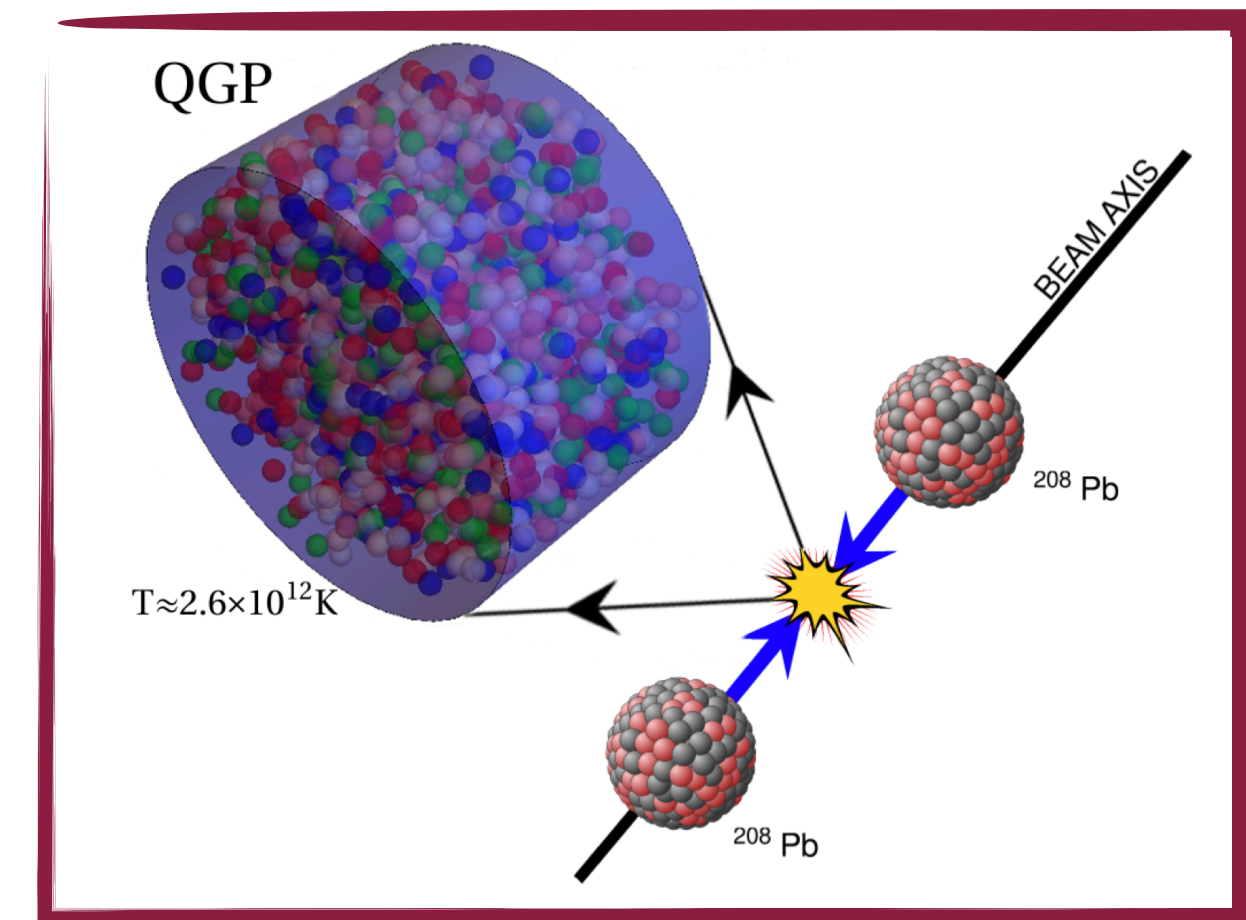


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# WAYS TO ACCESS THE PHASE DIAGRAM OF QCD

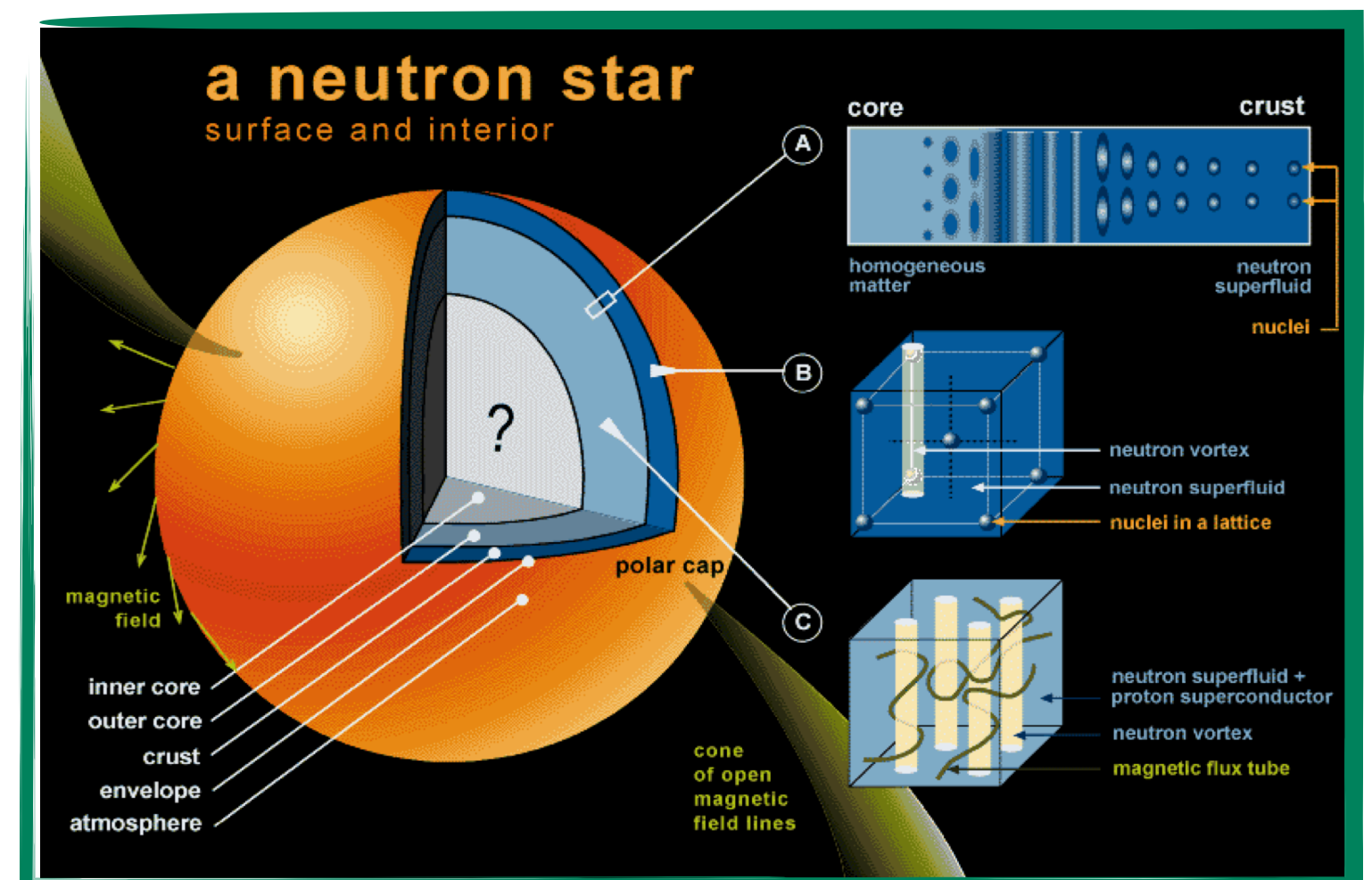


## Heavy-ion collision physics



*Nature Physics* 16, 615–619(2020)

## Neutron star physics



D.E. Á. Castillo, talk @RagTime 22

## lattice-QCD simulations

$$\mathcal{L} = \frac{1}{4g^2} G_{\mu\nu}^a G_{\mu\nu}^a + \sum_j \bar{q}_j (i\gamma^\mu D_\mu + m_j) q_j$$

where  $G_{\mu\nu}^a \equiv \partial_\mu A_\nu^a - \partial_\nu A_\mu^a + if_{bc}^a A_\mu^b A_\nu^c$

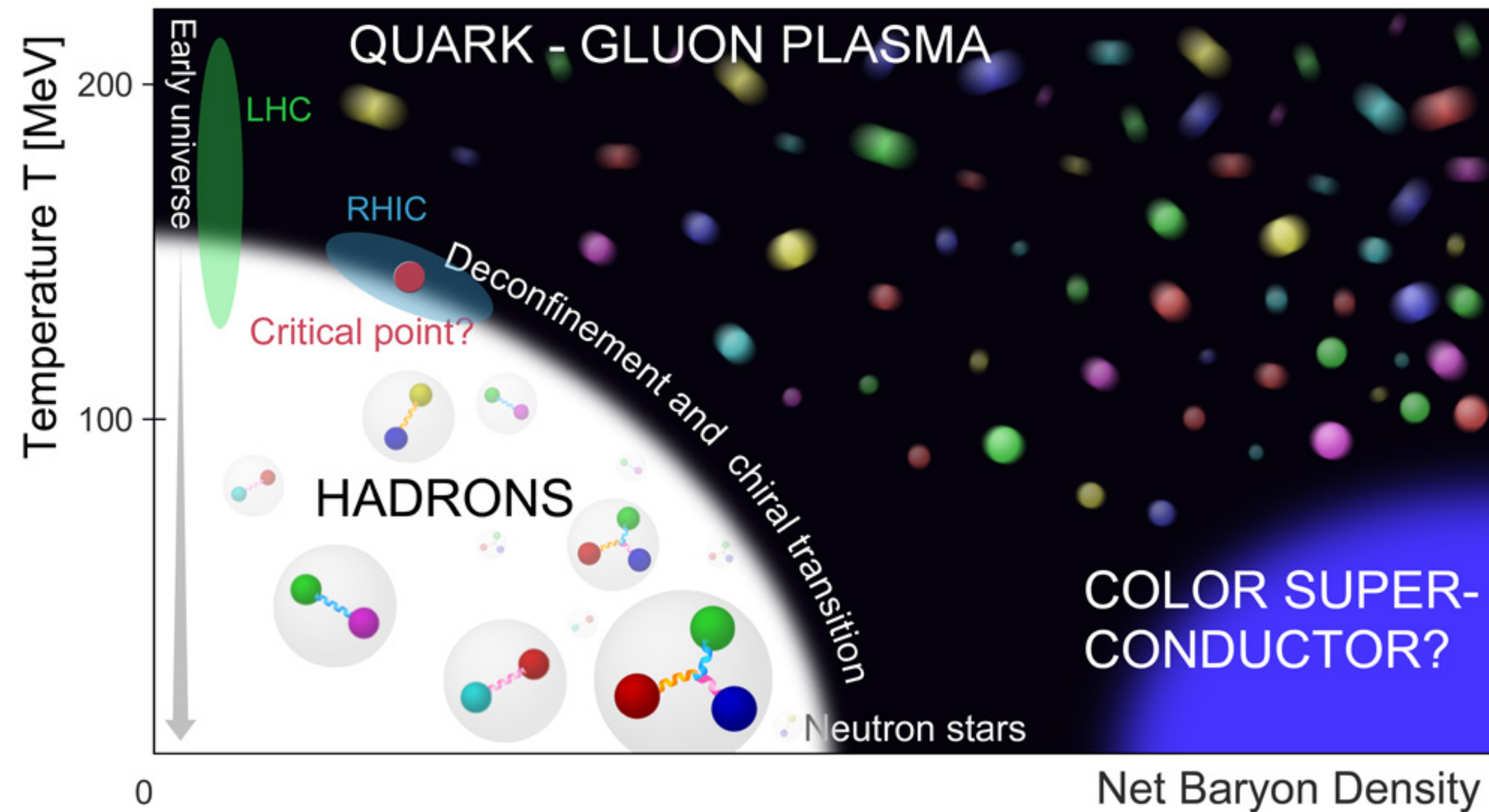
and  $D_\mu \equiv \partial_\mu + it^a A_\mu^a$

That's it!

D. Leinweber ([www.physics.adelaide.edu.au](http://www.physics.adelaide.edu.au))

# WHAT IS THE STRUCTURE OF QCD PHASE DIAGRAM?

Joint Institute for Computational Fundamental Science



## Vanishing $\mu_B$ , high $T$ (covered by lattice QCD)

- crossover transition
- no QCD critical point (CP) indicated by lattice QCD at  $\mu_B < 400$  MeV,  $T > 140$  MeV

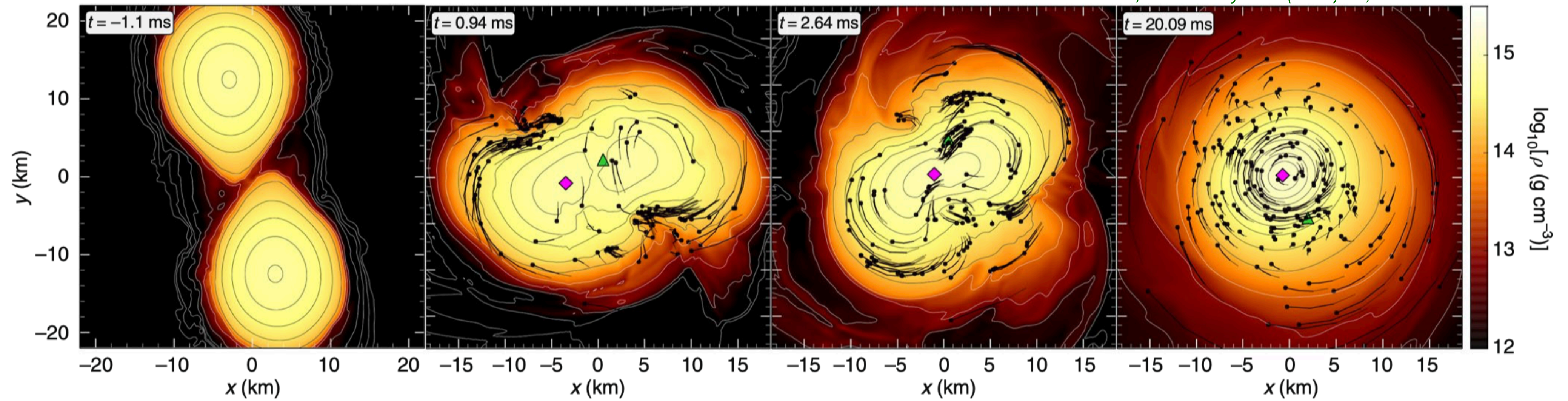
## Large $\mu_B$ , moderate $T$ (covered by QCD inspired models)

- 1st order transition (?)
- QCD CP (?)

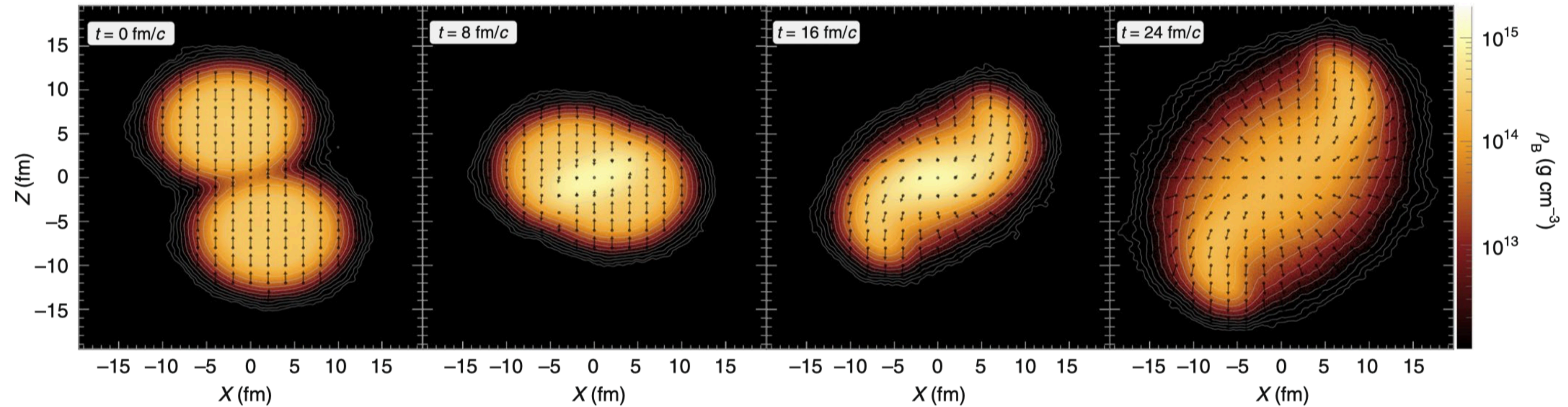
# NUCLEAR COLLISIONS ALLOW TO PROBE “WARM” DENSE MATTER

HADES, *Nature Phys.* 15 (2019) 10, 1040-1045

binary neutron  
star merger



noncentral Au+Au  
collision at 2.4 GeV  
per nucleon pair.



# MAPPING PHASE DIAGRAM WITH STATISTICAL HADRONIZATION MODEL

**Thermal models** of hadron production (based on the idea of **statistical hadronization**) have been very successful in **describing hadron yields in various collision processes**.

*J. Cleymans, H. Satz, F. Becattini, M. Gazdzicki, J. Sollfrank, W. Florkowski, W. Broniowski, J. Letessier, J. Rafelski, R. Stock, M. I. Gorenstein, A. Andronic, P. Braun-Munzinger, K. Redlich, and J. Stachel ...*

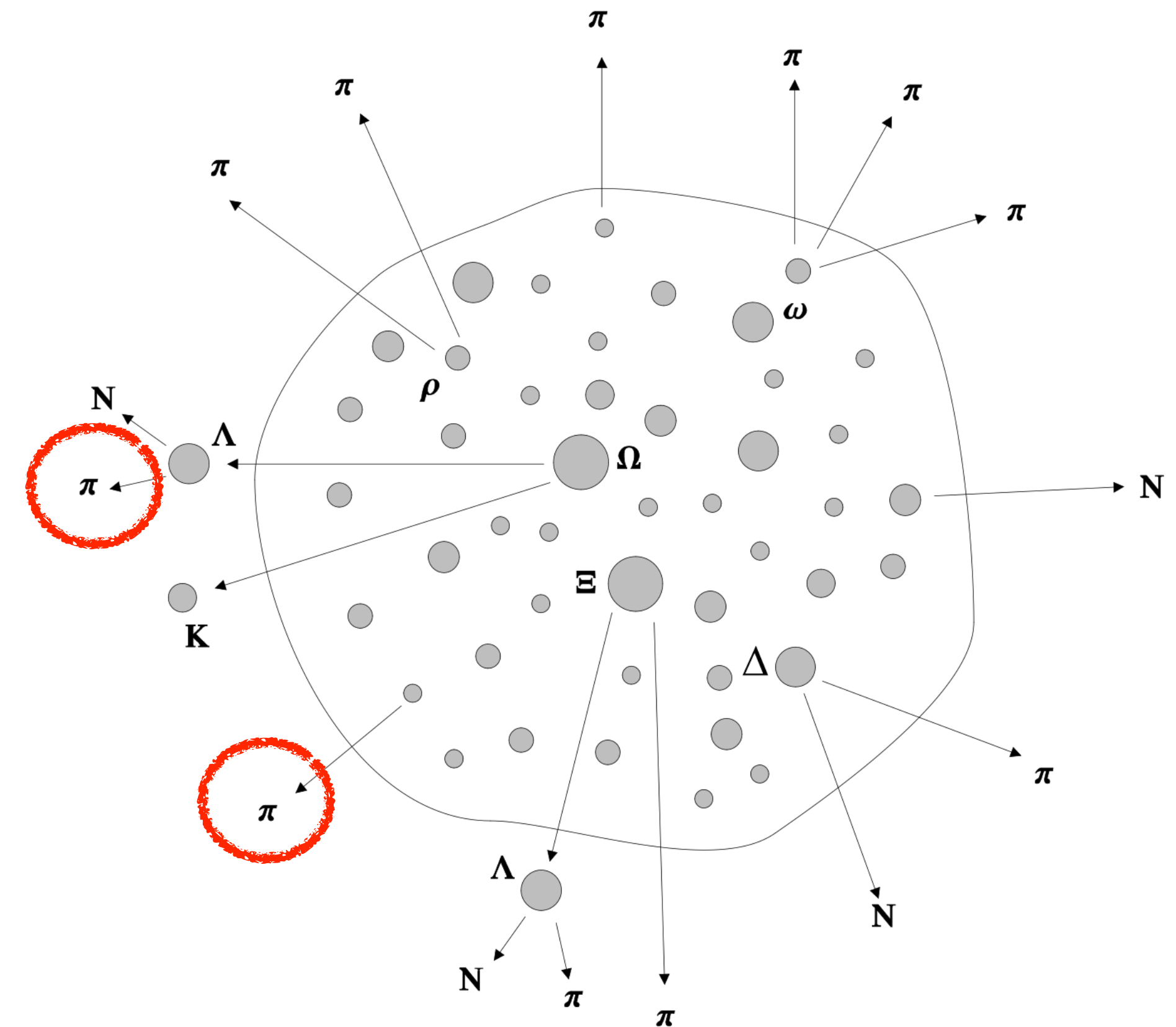
Matter formed at the chemical freeze-out is treated as multicomponent **non-interacting hadron resonance gas**.

In chemical equilibrium multiplicities of particle specie  $i$  can be written as:

$$N_i = V n_i = V g_i \int d^3p f_i(p)$$

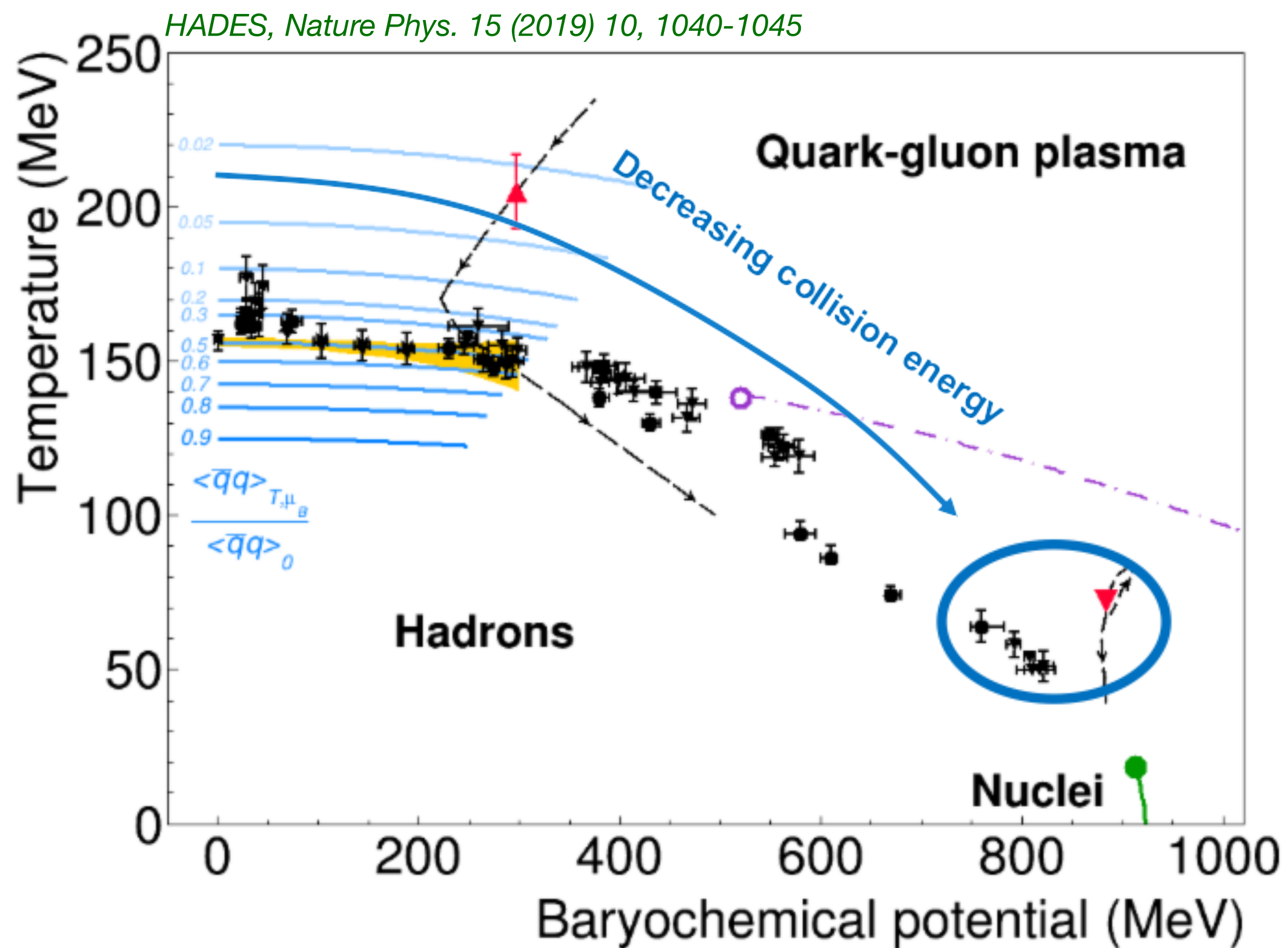
$$f_i(p) = \frac{1}{(2\pi)^3} \left[ \exp\left(\frac{E_i(p) - \mu_i}{T}\right) + \epsilon \right]^{-1}$$

One can **fit the ratios of measured particle yields** and **extract free parameters** giving location in the phase diagram.



*M. Michalec PhD thesis*

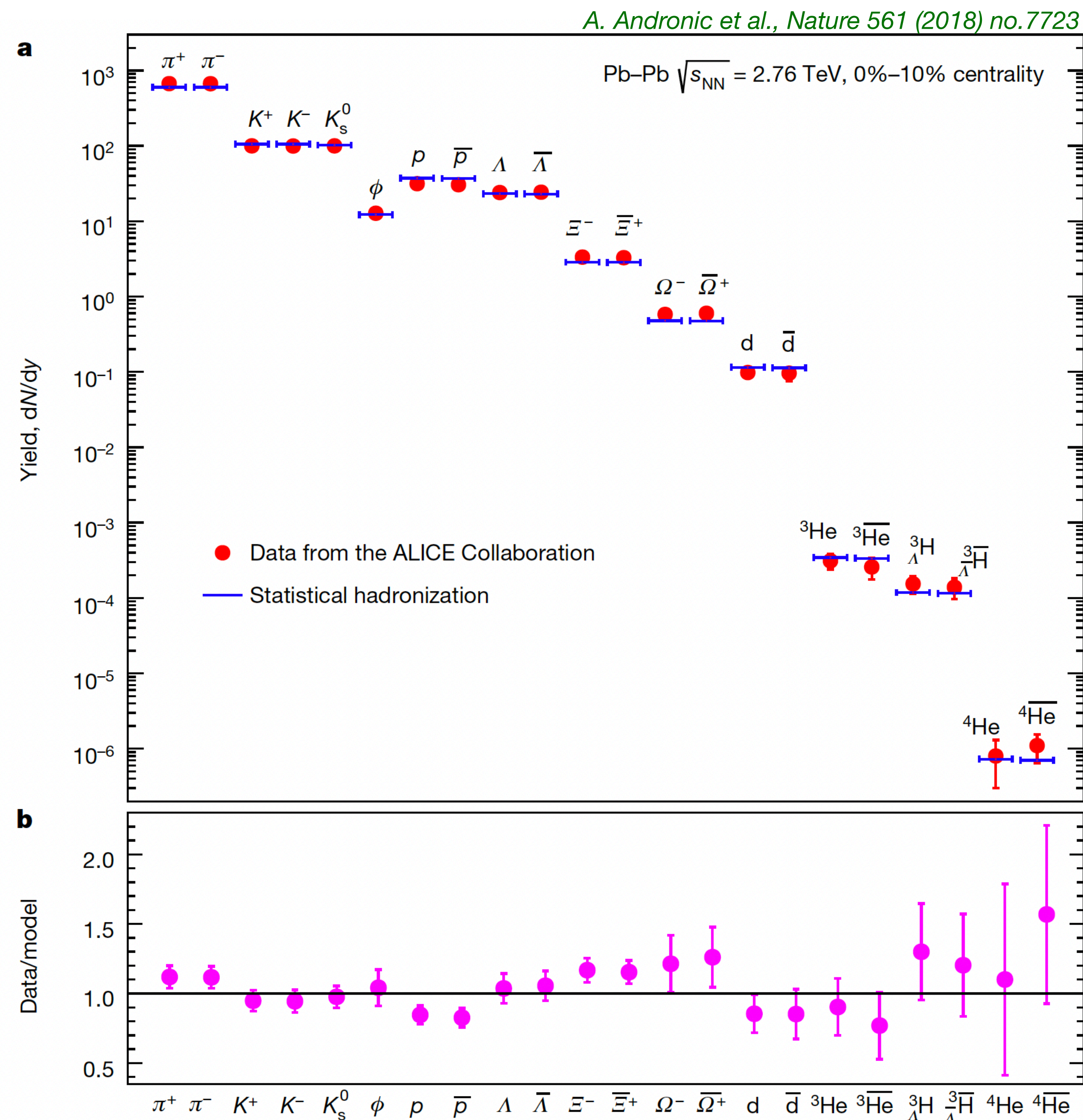
# MAPPING PHASE DIAGRAM WITH STATISTICAL HADRORIZATION MODEL



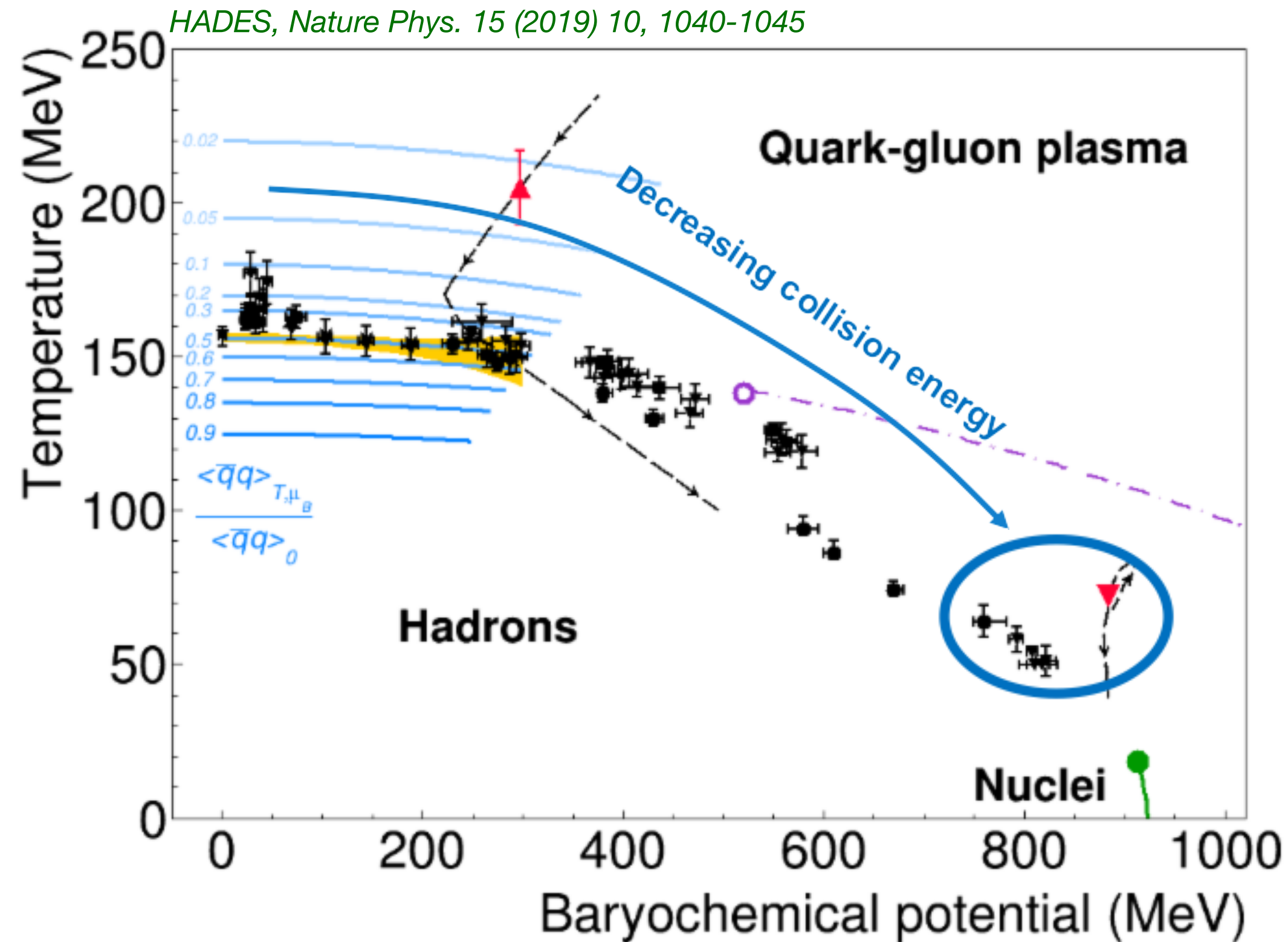
The **hadron yields** can be explained over several orders of multiplicity by fixing just **a few thermodynamic parameters**.

The role of hadronic resonances is crucial **at high energies** (~400 states are included).

At lower energies their role is diminished.



# MAPPING PHASE DIAGRAM WITH STATISTICAL HADRONIZATION MODEL



## Is it valid to assume equilibrium at low energies?

- Low number of newly produced particles in the interaction zone (~40 in central events, mainly pions)

## On the other hand:

- Original nucleons stopped in the interaction zone (~300 particles in central events)
- Longer life-time of the system – enough to thermalize

**The problem of whether the fireball in few-GeV energy regime is thermalized remains a matter of debate.**

The study of **hadron yields** and **spectra** is crucial to answer this question.

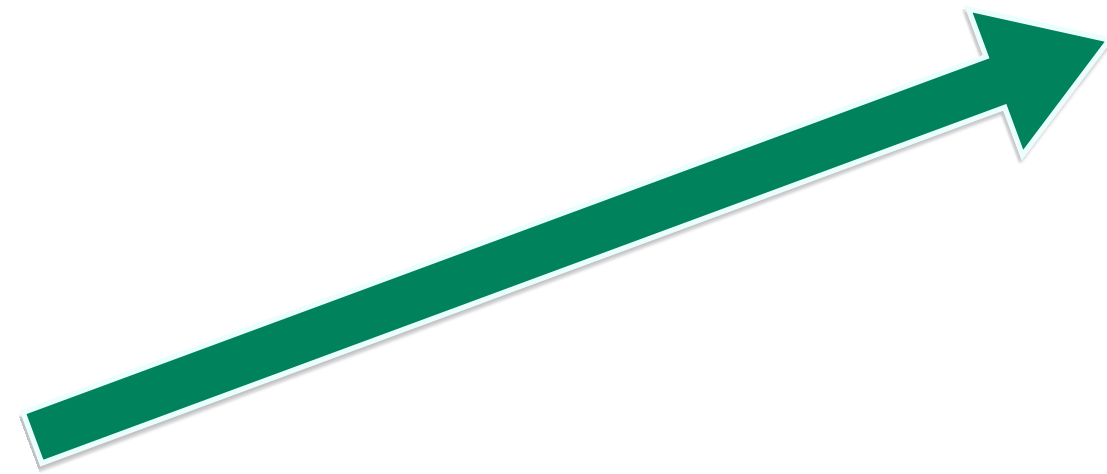
# DYNAMICAL MODELING OF HEAVY-ION COLLISIONS

## Standard prescription at high beam energies (RHIC/LHC):

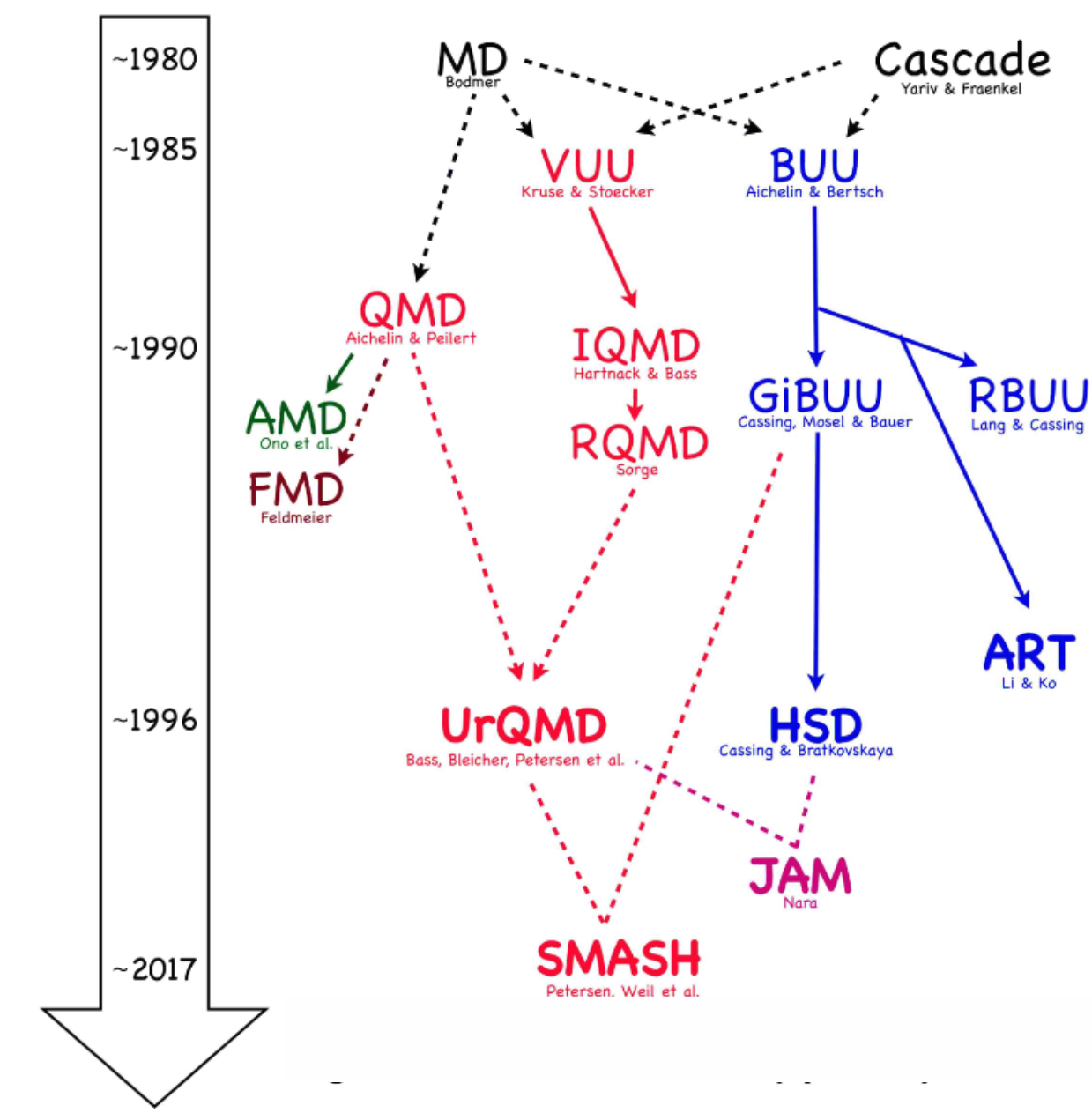
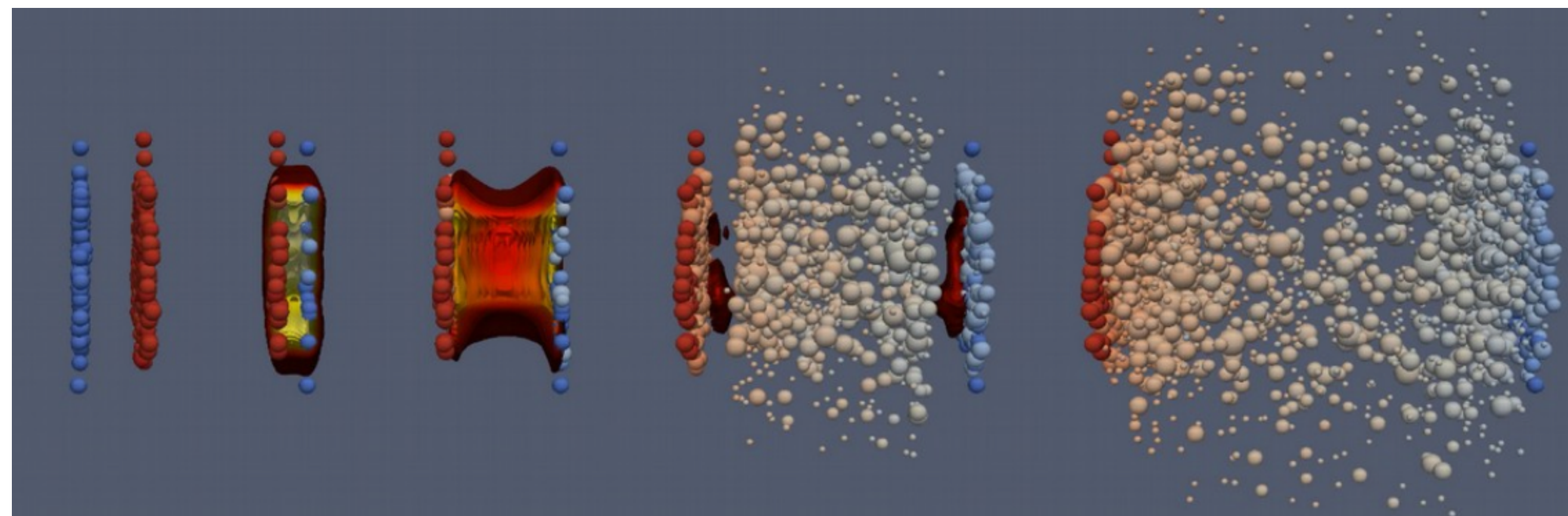
- Non-equilibrium initial conditions
- Viscous hydrodynamic evolution
- Equilibrium
- Hadronic final-state rescattering

## Standard prescription at low beam energies (GSI/FAIR):

- Hadronic transport
- Importance of:
  - Resonance dynamics
  - Nuclear potentials



MADAI collaboration, Hannah Petersen and Jonah Bernhard



Steffen A. Bass



# A FEW-GEV REGIME IS STILL PUZZLING FOR TRANSPORT MODELS

Basic description is obtained with transport models (UrQMD) and the emphasis is usually put on non-equilibrium features.

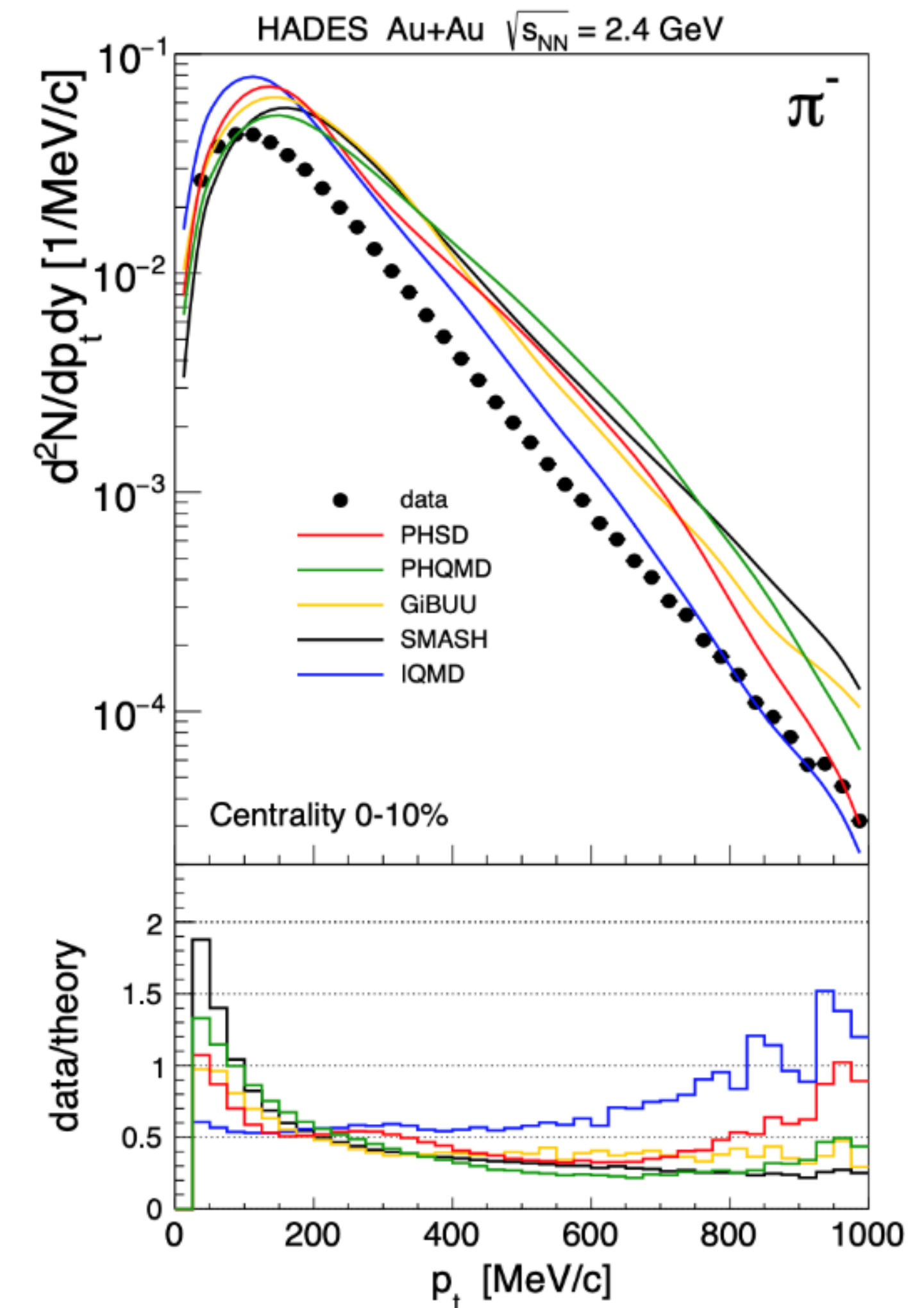
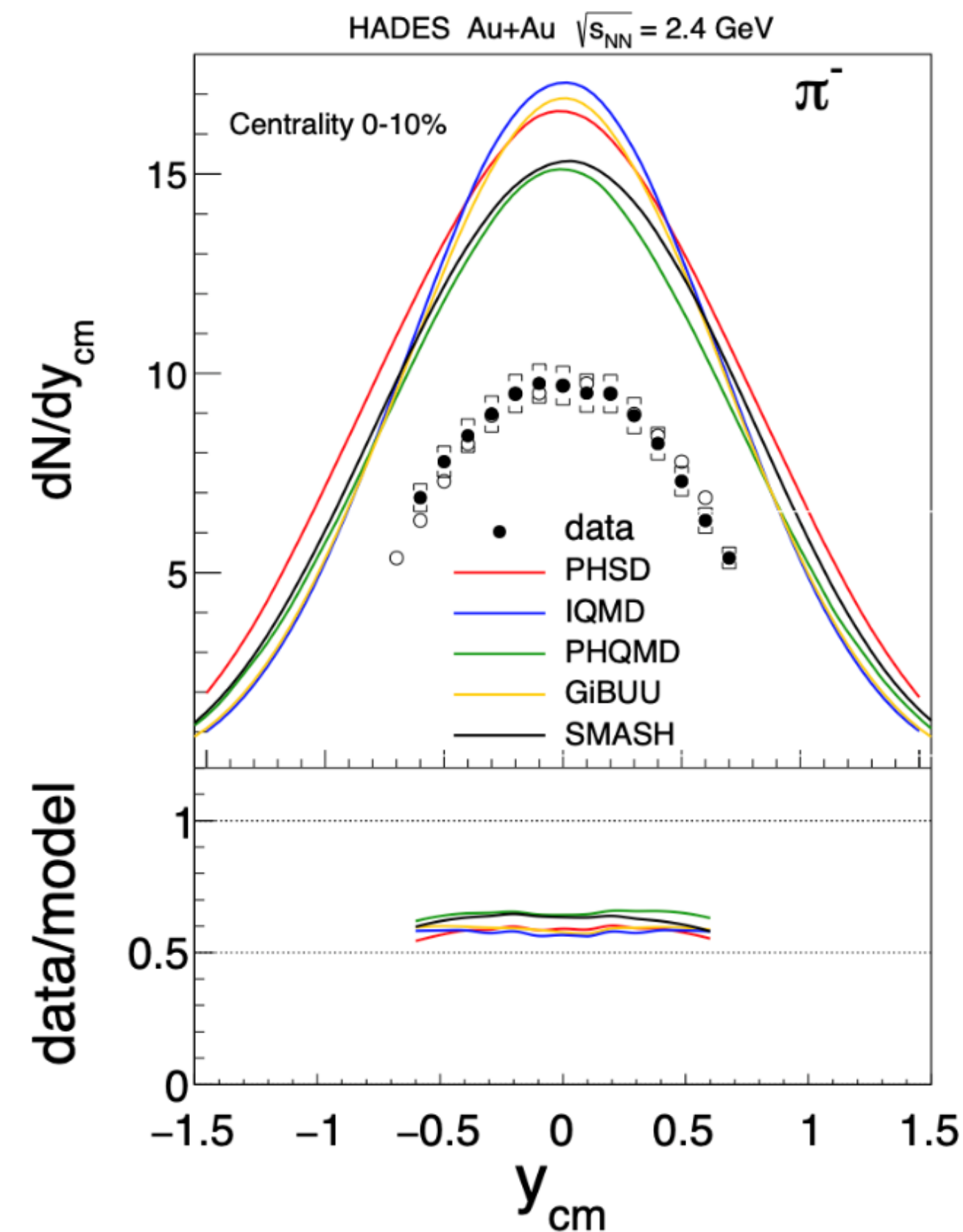
*S. A. Bass et al., Prog. Part. Nucl. Phys. 41, 255 (1998),*

*O. Buss, et al, Phys. Rept. 512, 1 (2012),*

*H. Petersen, D. Oliinychenko, M. Mayer, J. Staudenmaier, and S. Ryu, Nucl. Phys. A 982, 399 (2019),*

*C. Hartnack, R. K. Puri, J. Aichelin, J. Konopka, S. A. Bass, H. Stoecker, and W. Greiner, Eur. Phys. J. A 1, 151 (1998)*

*W. Cassing and E. L. Bratkovskaya, Phys. Rept. 308, 65 (1999).*



# HYDRO-INSPIRED MODELS

Instead of determining freeze-out conditions from hydrodynamic simulations or transport one can **model the freeze-out conditions (hypersurface and flow)**.

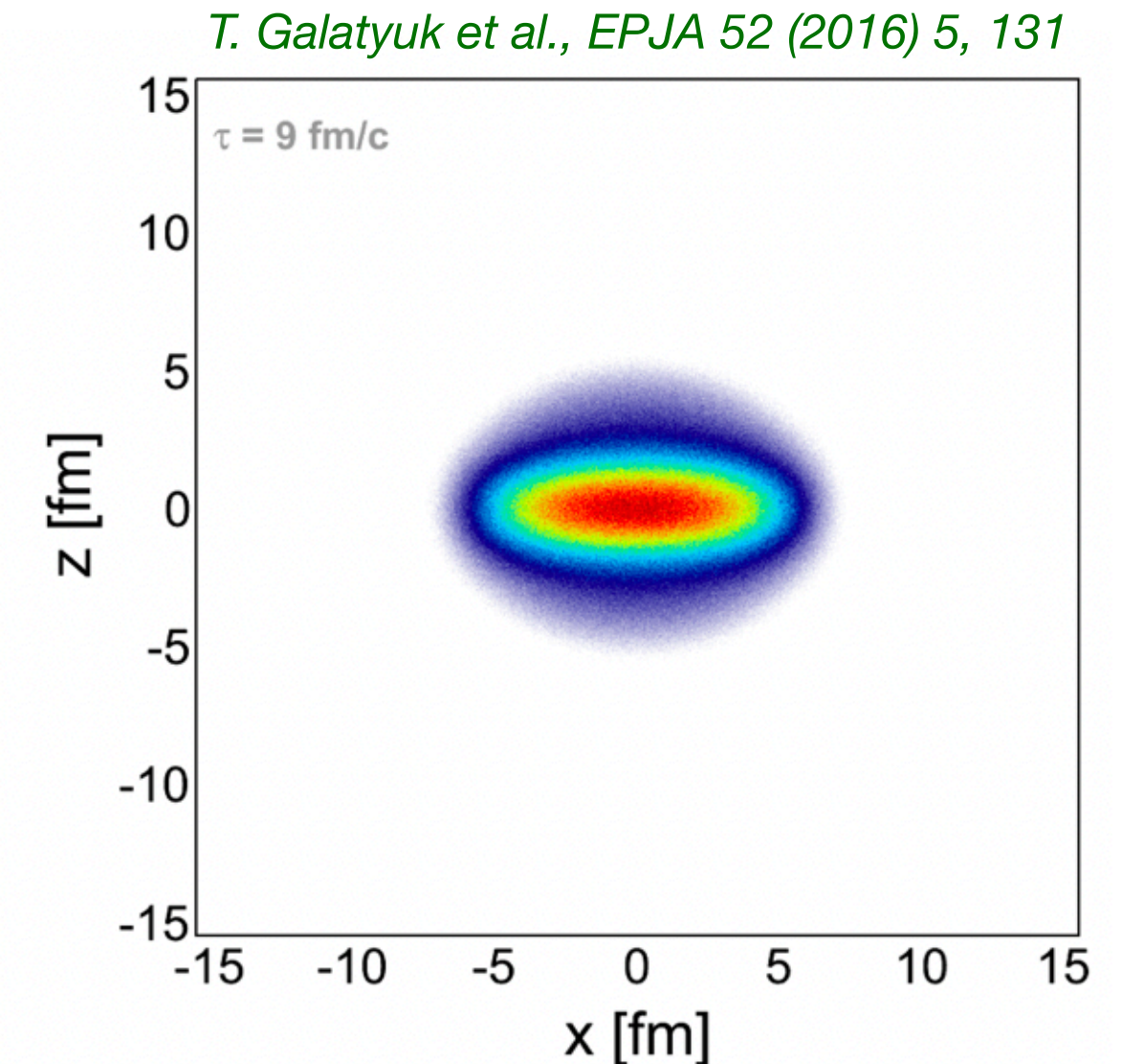
In the original formulation of the **blast-wave model by Siemens and Rasmussen (SR)** the **freeze-out was spherical and the flow was radial**.

*P. Siemens and O. Rasmussen, PRL 42, 880 (1979)*

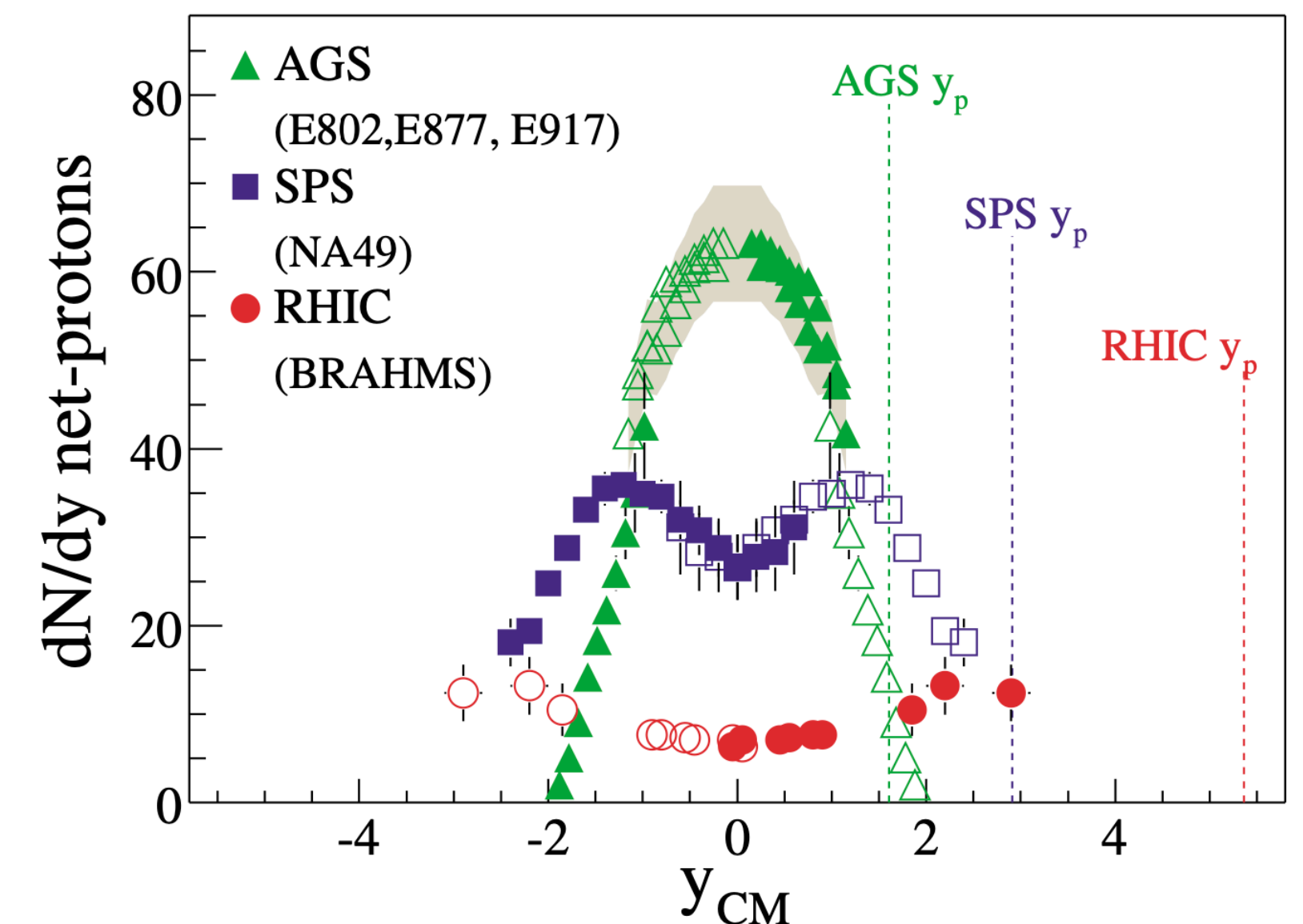
This approach was **modified for higher energies (RHIC and LHC) assuming boost-invariance and cylindrical symmetry** and used extensively by experimentalists.

*E. Schnedermann, J. Sollfrank, U. Heinz, PRC 48, 2462 (1993)*

**We aim to re-examine RS model in the context of low-energy collision measurements performed by HADES where boost-invariance is not observed.**



*I. G. Bearden et al. (BRAHMS), PRL 93, 102301 (2004)*



# COOPER-FRYE FORMULA

**Invariant momentum spectrum** of particles emitted from an expanding source through the hypersurface  $\Sigma_\mu$

*F. Cooper and G. Frye, PRD 10, 186 (1974).*

$$E_p \frac{dN}{d^3p} = \int d^3\Sigma_\mu(x) p^\mu f(x, p)$$

$$E_p = \sqrt{m^2 + \mathbf{p}^2}.$$

Assuming a **spherically symmetric source** the freeze-out points are defined by the space-time coordinates

$$x^\mu = (t, \mathbf{x}) = (t(\zeta), r(\zeta) \mathbf{e}_r)$$

$$\mathbf{e}_r = (\cos \phi \sin \theta, \sin \phi \sin \theta, \cos \theta) \quad \zeta \longrightarrow (t(\zeta), r(\zeta))$$

$$d^3\Sigma_\mu = -\epsilon_{\mu\alpha\beta\gamma} \frac{\partial x^\alpha}{\partial a} \frac{\partial x^\beta}{\partial b} \frac{\partial x^\gamma}{\partial c} da db dc.$$

$$d^3\Sigma_\mu = (r'(\zeta), t'(\zeta) \mathbf{e}_r) r^2(\zeta) \sin \theta d\theta d\phi d\zeta.$$

We assume **sudden freezeout**

$$t(r) = \text{const}$$

With the hadron four-momentum parametrized as  $p^\mu = (E_p, p \mathbf{e}_p)$   $\mathbf{e}_p = (\cos \varphi \sin \vartheta, \sin \varphi \sin \vartheta, \cos \vartheta)$

We get

$$d^3\Sigma(x) \cdot p = E_p \sin \theta d\theta d\phi r^2 dr$$

# LOCAL THERMAL EQUILIBRIUM

Assume that the hadron system formed is very close to **local thermodynamic equilibrium**

$$f(x, p) = \frac{g_s}{(2\pi)^3} \left[ \Upsilon^{-1} \exp\left(\frac{p \cdot u}{T}\right) - \chi \right]^{-1}$$

The fugacity is defined as

*G. Torrieri, S. Steinke, W. Broniowski, W. Florkowski, J. Letessier, and J. Rafelski, CPC 167, 229 (2005).*

$$\Upsilon = \gamma_q^{N_q + N_{\bar{q}}} \gamma_s^{N_s + N_{\bar{s}}} \exp\left(\frac{\mu}{T}\right) \quad \mu = \sum_Q Q \mu_Q \quad Q \in \{B, I_3, S\}$$

We allow for **strangeness undersaturation** (characteristic feature at low beam energies).

*J. Rafelski, J. Letessier, and A. Tounsi, Acta Phys. Pol. B28 (1997) 2841*

# HUBBLE-LIKE RADIAL FLOW

We introduce a **spherically symmetric flow**

$$u^\mu = \gamma(r)(1, v(r)e_r)$$

In the **original SR blast-wave model**, it was assumed that the thermodynamic parameters as well as the radial flow velocity are constant

$$(T = \text{const}, \mu = \text{const}, v = v_0 = \text{const})$$

We take **Hubble-like flow**

$$v(r) = \tanh(Hr)$$

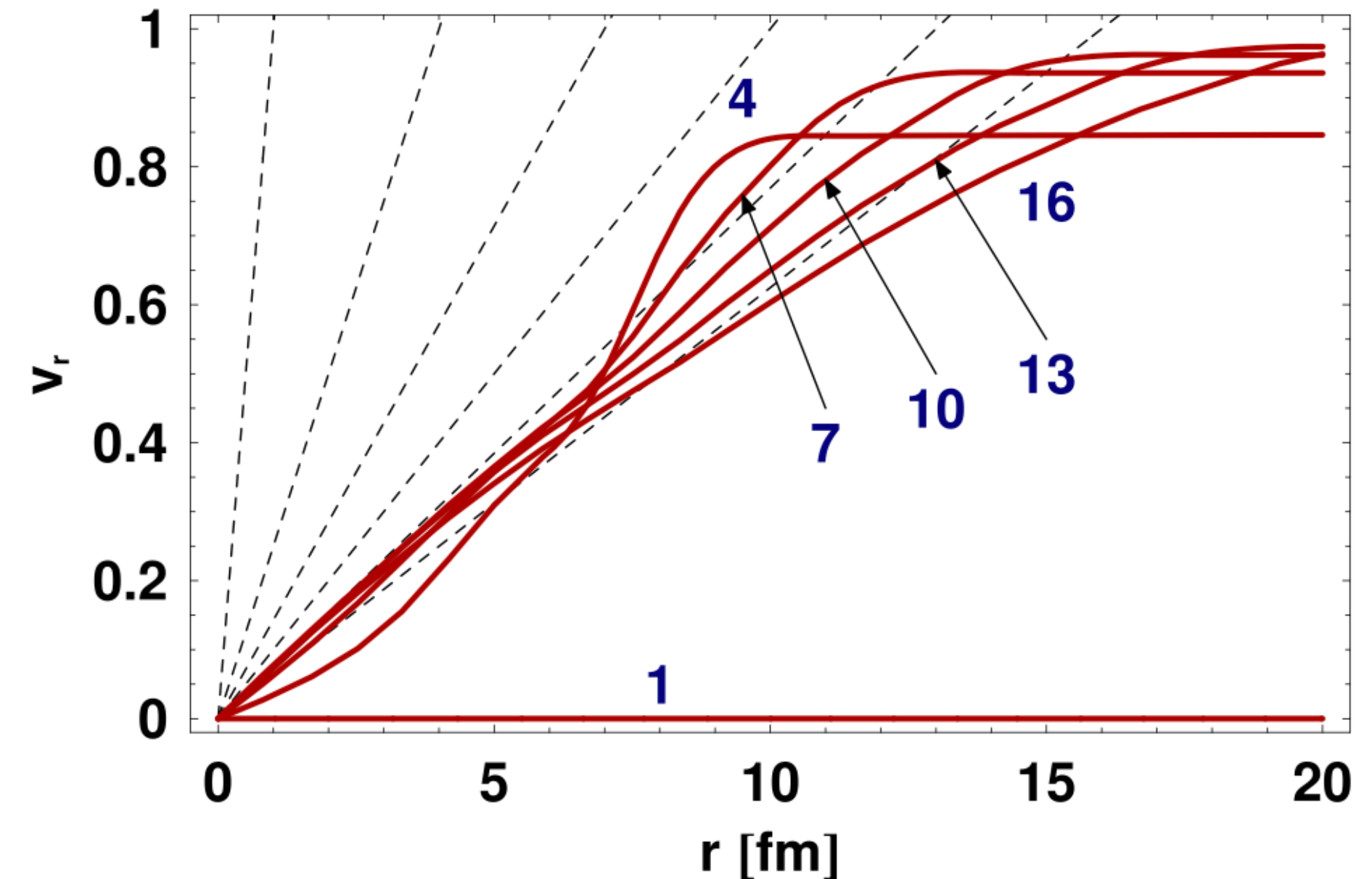
The parameter **H** plays a role of the **Hubble constant** in the theory of expanding Universe.

As a result we get

$$p \cdot u = \gamma(E_p - pv\kappa)$$

$$\kappa \equiv e_p \cdot e_r$$

*M. Chojnacki, W. Florkowski, and T. Csorgo, PRC 71, 044902 (2005).*



Condition of constant radial flow breaks requirement that the flow at the center of the system should vanish.

Results of hydrodynamic calculations indicate that the **radial flow linearly grows with radius** for small values of  $r$ .

# THERMINATOR

Our freeze-out prescription is implemented in the **THERMINATOR** Monte Carlo hadron generator which allows for studies of hadron production taking place on **arbitrary freeze-out hypersurfaces** defined in the four-dimensional space-time.

*A. Kisiel, T. Taluc, W. Broniowski, and W. Florkowski, Comput. Phys. Commun. 174, 669 (2006).*

*M. Chojnacki, A. Kisiel, W. Florkowski, and W. Broniowski, Comput. Phys. Commun. 183, 746 (2012).*

THERMINATOR generates primordial particles at the freeze-out hypersurface (~400 states are included).

Unstable particles are free-streaming and are allowed to decay contributing to feed-down.

The code includes contributions from decays of all heavier resonances — most of them are very small or negligible.

The largest contribution comes from decays of the lowest-lying baryonic resonance, i.e., Delta (1232).

# THERMODYNAMIC PARAMETERS

**We obtain thermodynamic model parameters from the ratios of experimental yields measured by HADES in Au+Au collisions at 2.4 GeV.**

We assume that the **protons finally bound in the emitted deuterons, tritons, and Helium nuclei were initially frozen out as unbound nucleons**; hence, they are included in the proton yield

**In the studies of the ratios of hadronic yields the invariant volume cancels (if the thermodynamic parameters are constant on the freeze-out hypersurface!).**

$$N = \int d^3\Sigma_\mu(x) \int \frac{d^3p}{E_p} p^\mu f(x, p).$$

$$N = n(T, \Upsilon) \int d^3\Sigma_\mu(x) u^\mu(x) \equiv n(T, \Upsilon) \mathcal{V},$$

TABLE I. Particle multiplicities used in the determination of the freeze-out parameters. Protons bound in nuclei are taken into account as shown.

Particle	Multiplicity	Uncertainty	Ref.
p	77.6	$\pm 2.4$	[29,31]
p (bound)	46.5	$\pm 1.5$	[29,31]
$\pi^+$	9.3	$\pm 0.6$	[32]
$\pi^-$	17.1	$\pm 1.1$	[32]
$K^+$	$5.98 \cdot 10^{-2}$	$\pm 6.79 \cdot 10^{-3}$	[33]
$K^-$	$5.6 \cdot 10^{-4}$	$\pm 5.96 \cdot 10^{-5}$	[33]
$\Lambda$	$8.22 \cdot 10^{-2}$	$^{+5.2}_{-9.2} \cdot 10^{-3}$	[34]

$\sqrt{s_{NN}} = 2.4$  GeV full phase space for the 10% Au-Au collisions

[29] M. Szala (HADES), *Light nuclei formation in heavy ion collisions measured with HADES*

[31] M. Szala (HADES), *Springer Proc.Phys.* 250 (2020) 297-301

[32] J. Adamczewski-Musch et al., (HADES) *EPJA* 56 (2020) 10, 259

[33] J. Adamczewski-Musch et al. (HADES), *PLB* 778, 403 (2018).

[34] J. Adamczewski-Musch et al. (HADES), *PLB* 793, 457 (2019).

# THERMODYNAMIC PARAMETERS

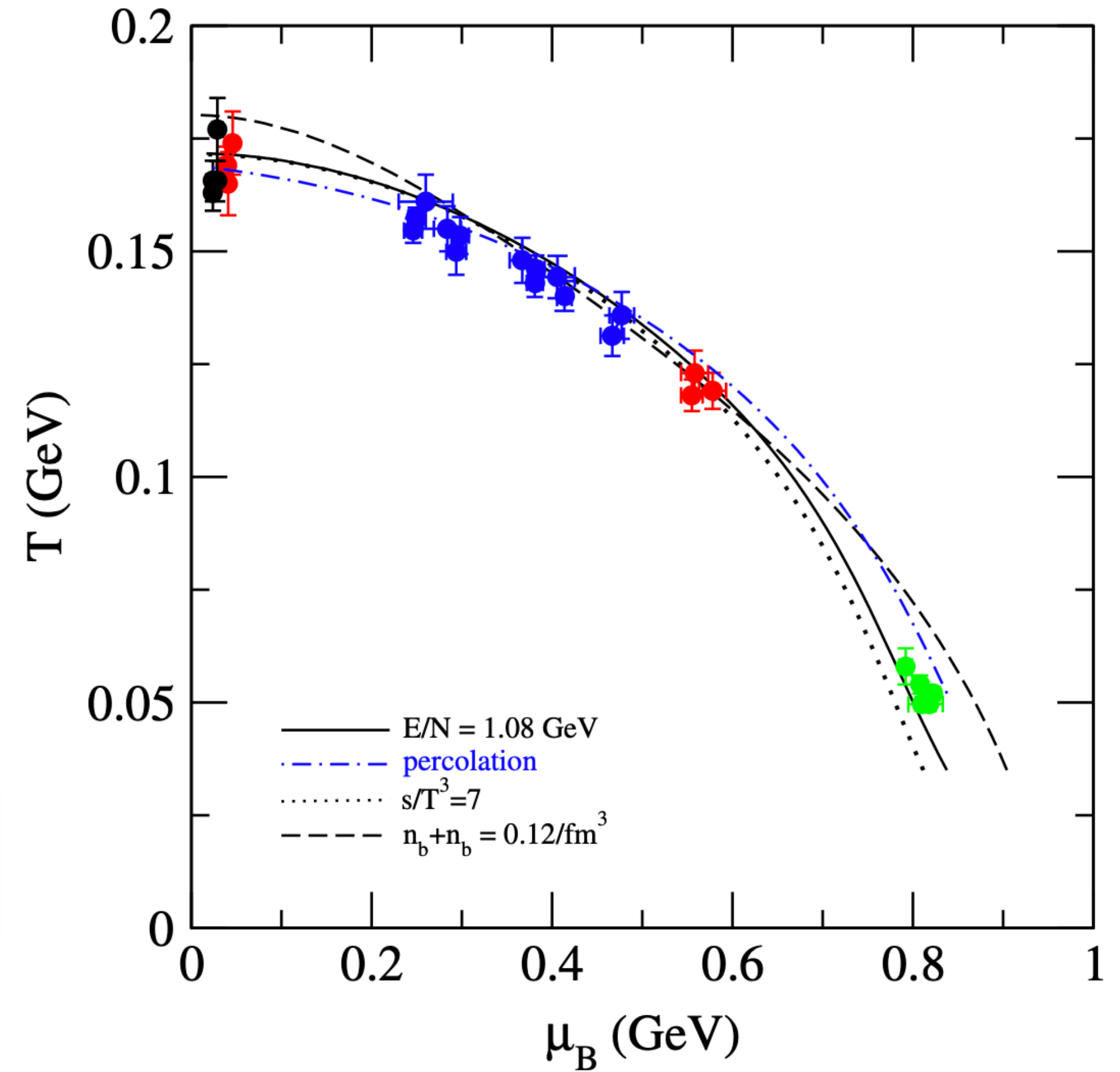
$$T = 49.6 \pm 1 \text{ MeV}, \mu_B = 776 \pm 3 \text{ MeV},$$

$$\mu_{I_3} = -14.1 \pm 0.2 \text{ MeV}, \mu_S = 123.4 \pm 2 \text{ MeV},$$

$$\gamma_s = 0.16 \pm 0.02$$

*J. Cleymans, H. Oeschler, K. Redlich, and S. Wheaton, PRC 73, 034905 (2006).*  
*P. Castorina, A. Iorio, D. Lanteri, H. Satz, and M. Spousta, PRC 101, 054902 (2020).*

*J. Cleymans, H. Oeschler, K. Redlich, and S. Wheaton, PRC 73, 034905 (2006).*





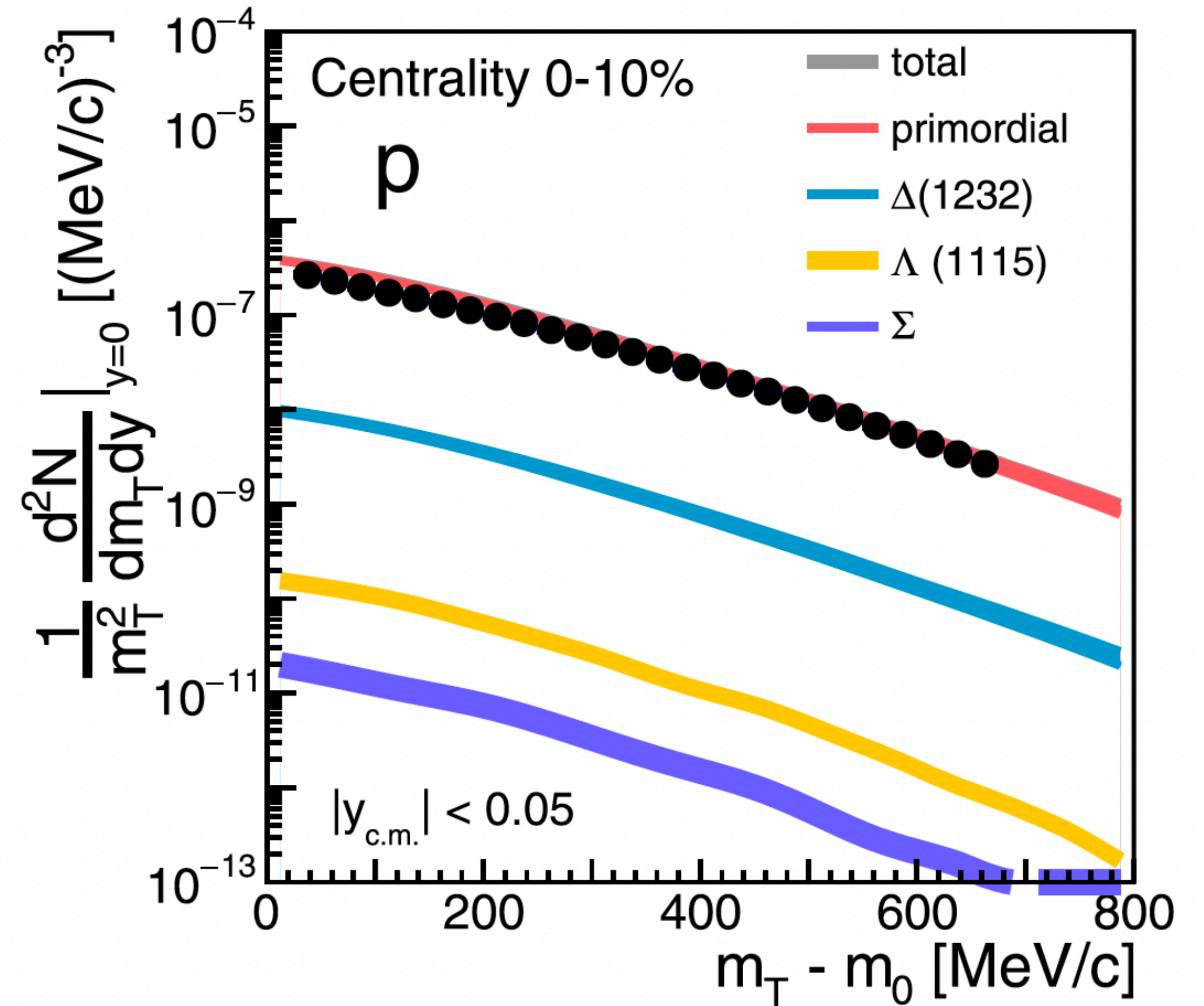
# TRANSVERSE-MOMENTUM AND RAPIDITY SPECTRA

For a fixed value of  $\mathbf{H}$ , the absolute normalization of the yields determines the value of  $\mathbf{R}$ . Hence we may treat  $\mathbf{R} = \mathbf{R}(\mathbf{H})$  and we are left with only one independent parameter  $\mathbf{H}$ .

Value of  $\mathbf{H}$  is obtained from the fit of the proton transverse-mass spectrum by minimization of the quadratic deviation

$$Q^2(H) = \sum_i \frac{(Q_{i,\text{model}}(H) - Q_{i,\text{exp}})^2}{Q_{i,\text{exp}}^2}$$

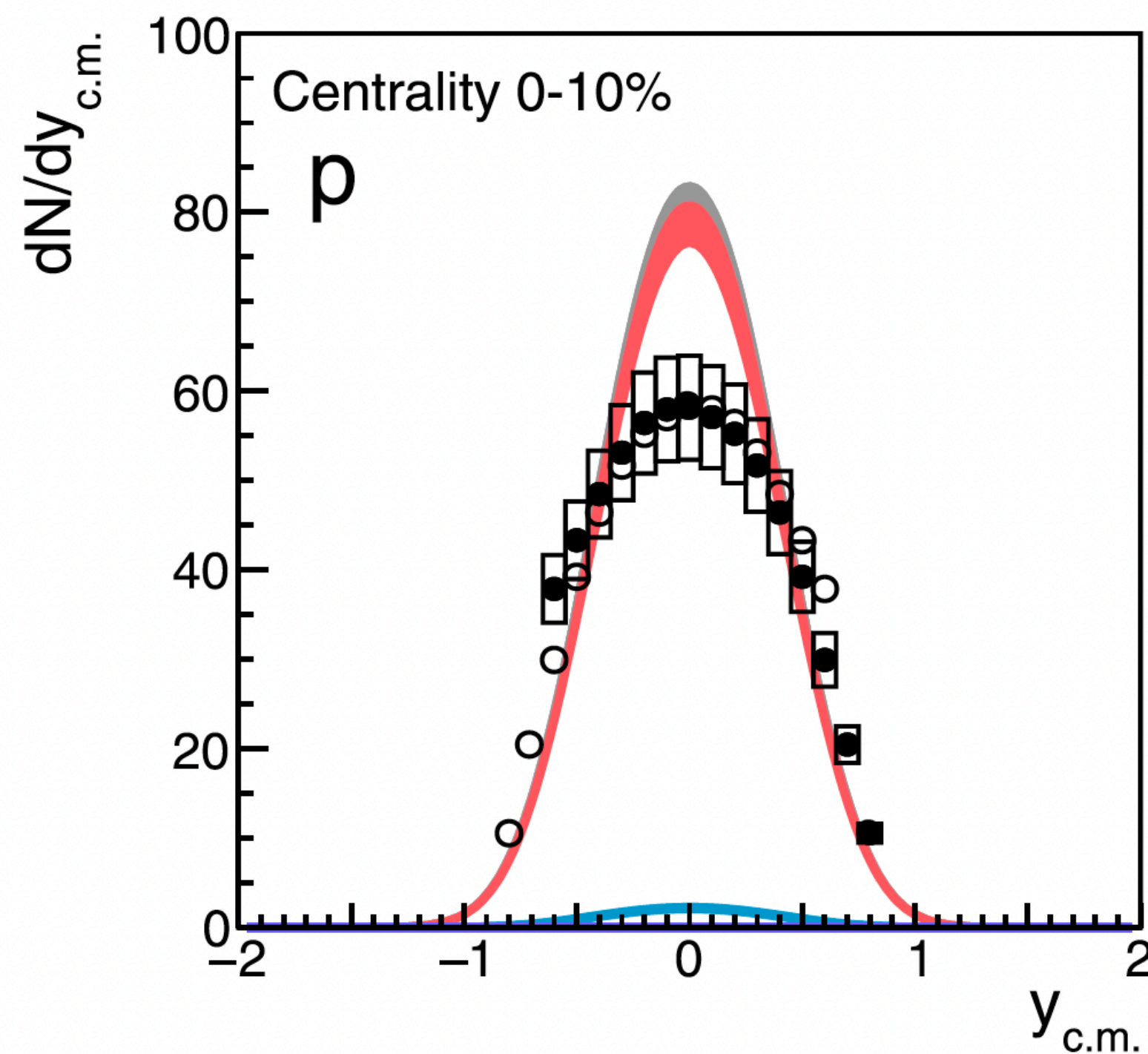
$$\begin{aligned} R &= 16.02 \text{ fm} \\ H &= 0.04 \text{ 1/fm} \end{aligned}$$



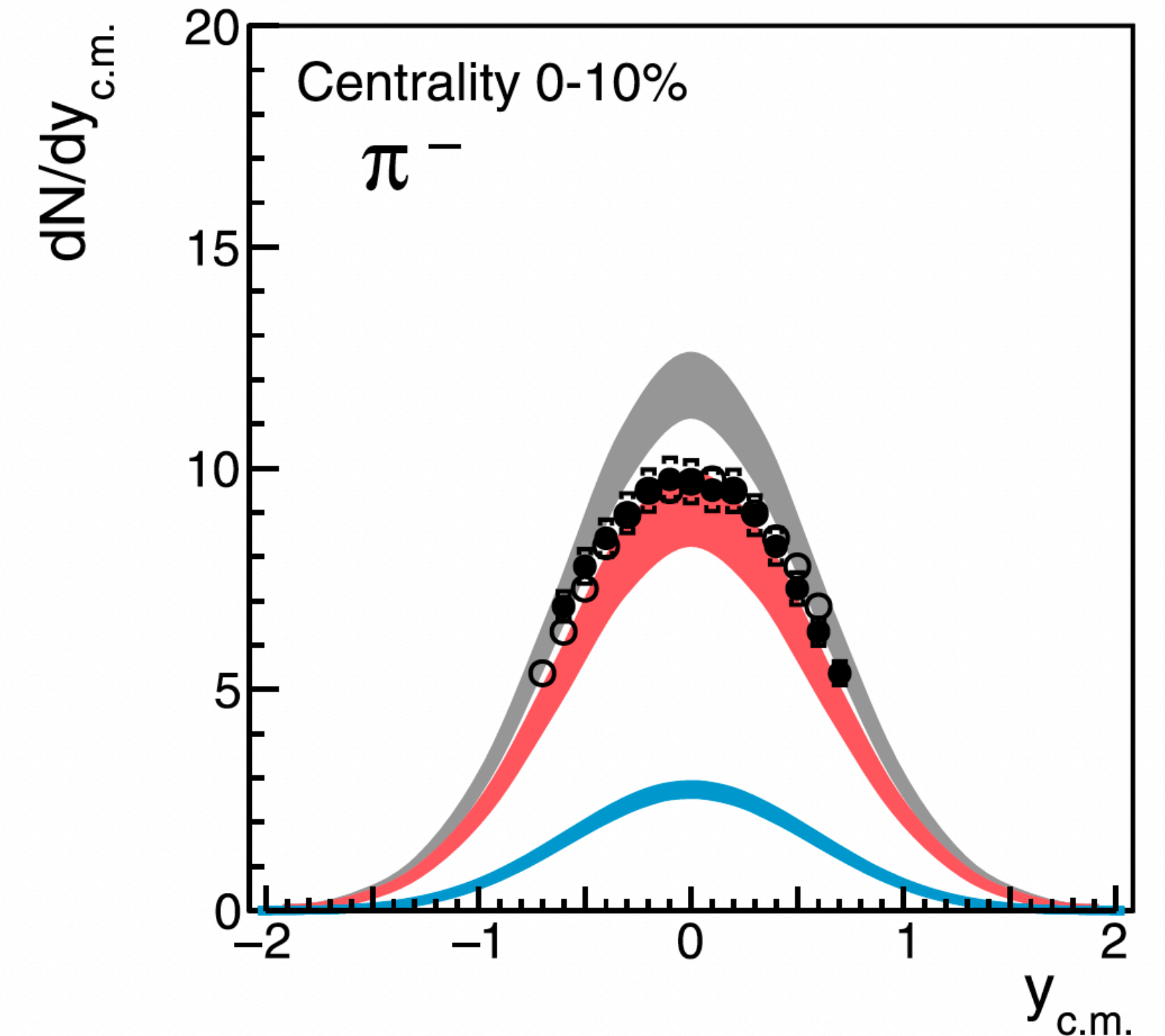
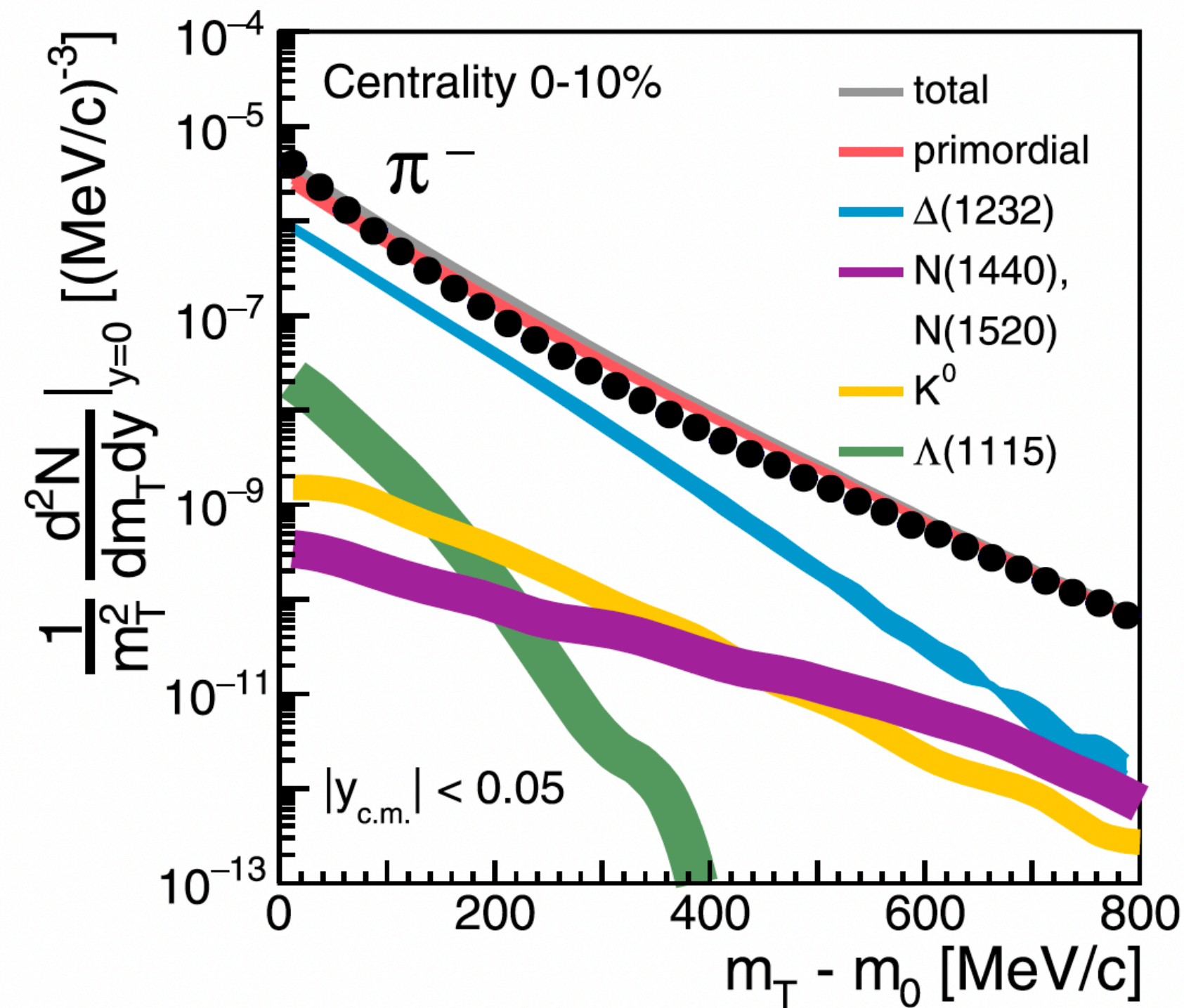
$$(Q = 0.20)$$

# TRANSVERSE-MOMENTUM AND RAPIDITY SPECTRA

Having determined the value of  $H$ , we can predict other model spectra.



$$Q = 0.28$$



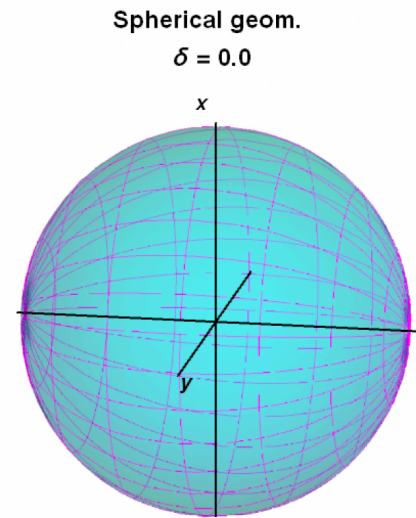
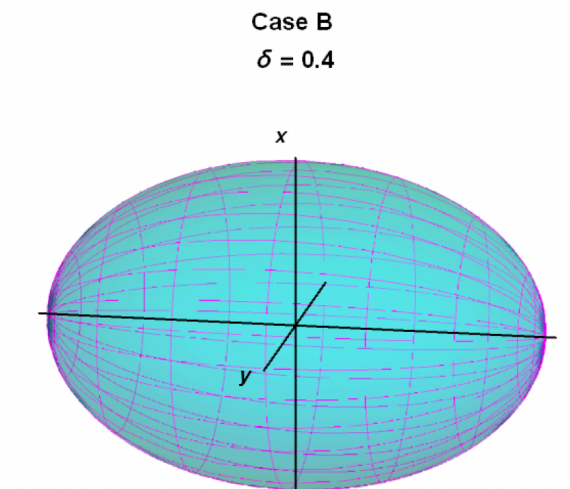
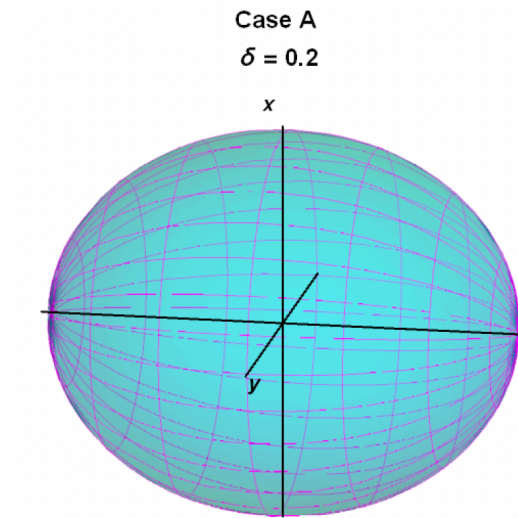
The fact that the rapidity distribution is equally well described (compared to the transverse-mass distribution) points out the **approximate (!) spherical symmetry** of the produced system.

# IMPROVEMENT: SPHEROIDAL EXPANSION

$$x^\mu = (t, r\sqrt{1-\epsilon}\sin\theta\hat{e}_\rho, r\sqrt{1+\epsilon}\cos\theta)$$

$$u^\mu = \gamma(\zeta, \theta) \left( 1, v(\zeta)\sqrt{1-\delta}\sin\theta\hat{e}_\rho, v(\zeta)\sqrt{1+\delta}\cos\theta \right)$$

$$\hat{e}_\rho = (\cos\phi, \sin\phi)$$

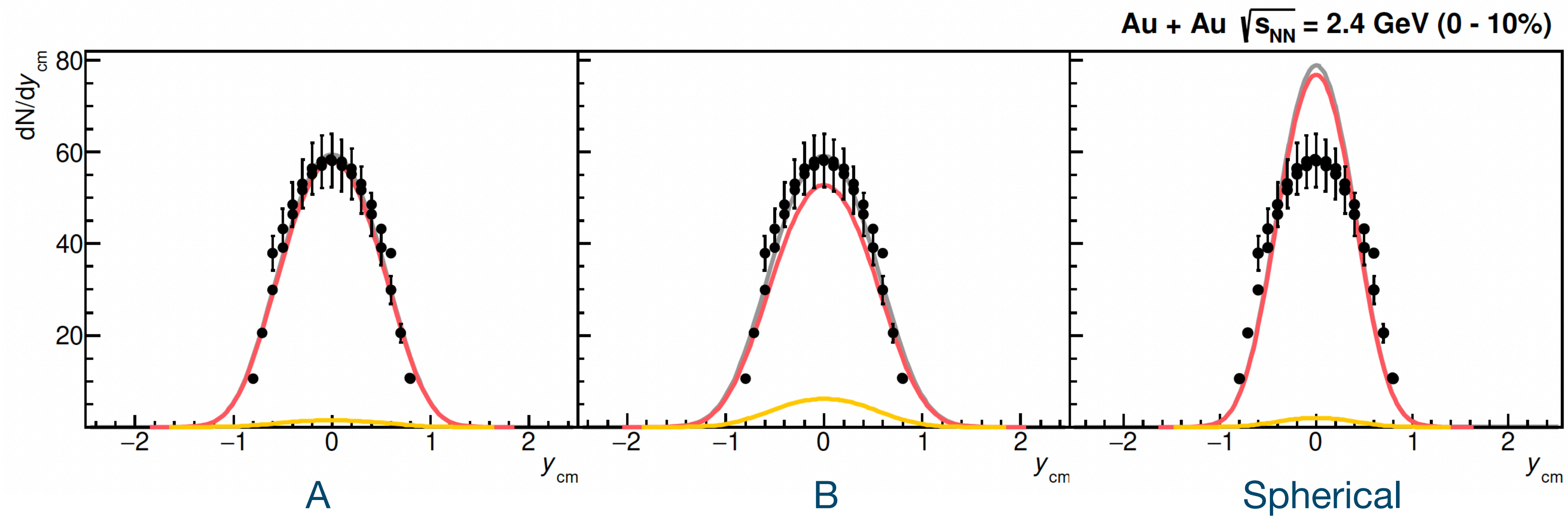
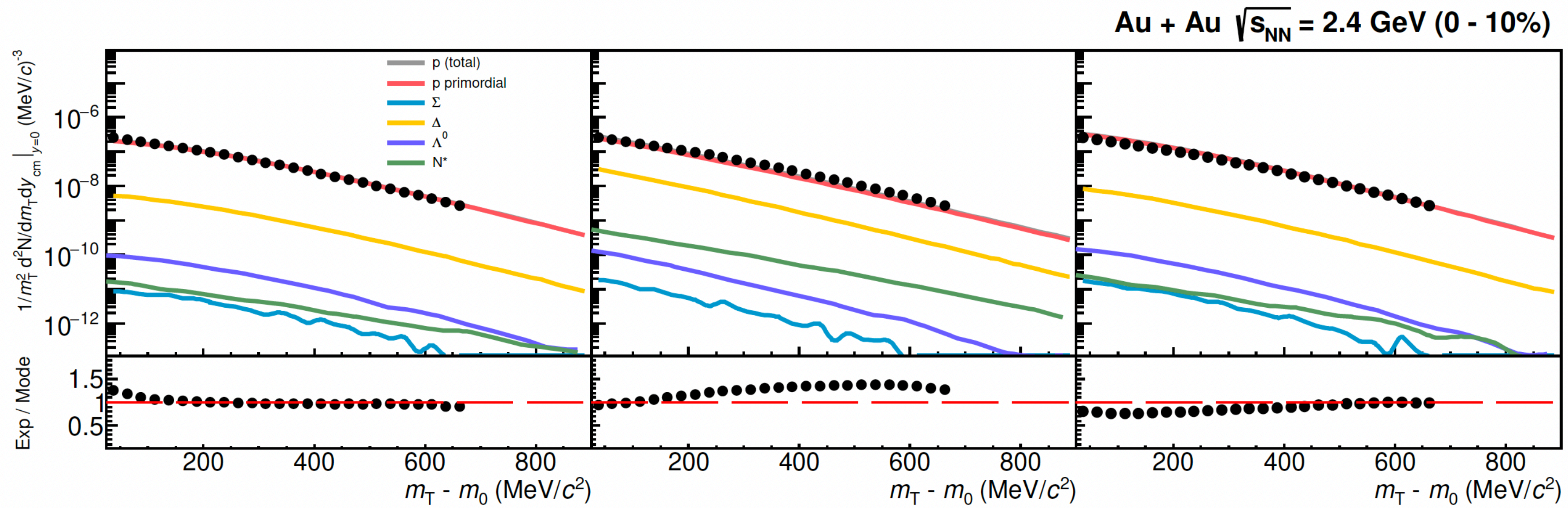


Parameter	Spherical geometry, Ref. [29]	Case A	Case B
$T$ (MeV)	49.6	49.6	70.3
$R$ (fm)	16.0	15.7	6.06
$\mu_B$ (MeV)	776	776	876
$\mu_S$ (MeV)	123.4	123.4	198.3
$\mu_{I_3}$ (MeV)	-14.1	-14.1	-21.5
$\gamma_S$	0.16	0.16	0.05
$\chi^2/N_{df}$	$N_{df} = 0$	$N_{df} = 0$	1.13/2
$H$ (GeV)	0.008	0.01	0.0225
$\delta$	0	0.2	0.4
$\sqrt{Q^2}$	0.285	0.238	0.256

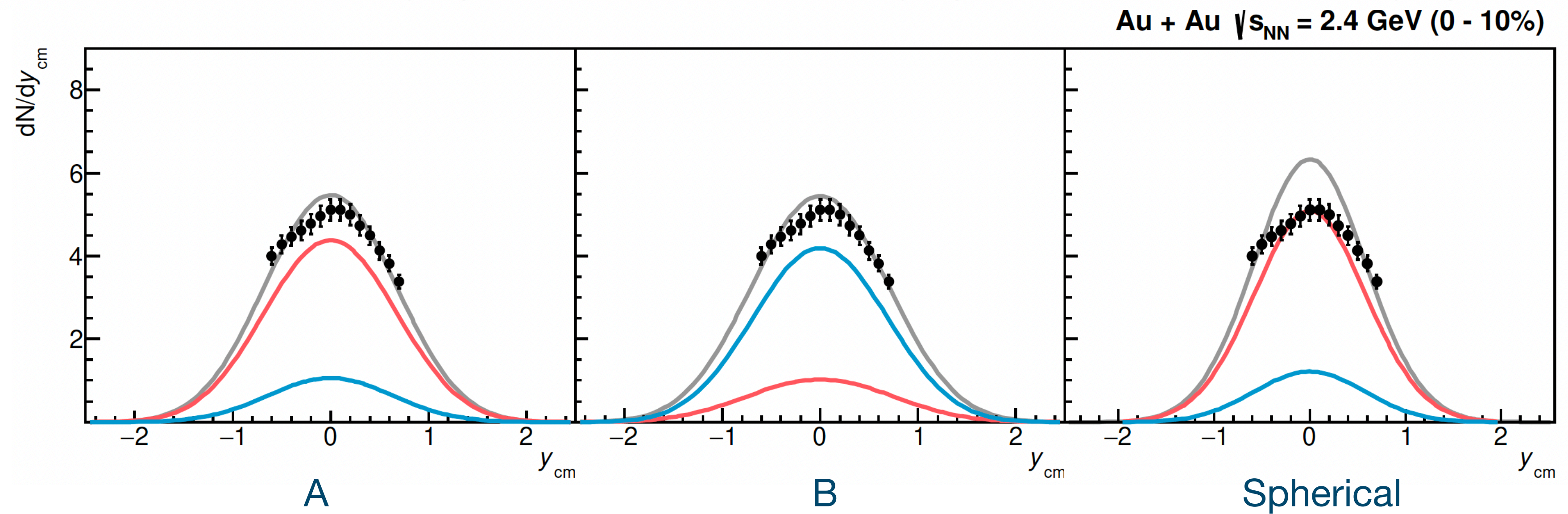
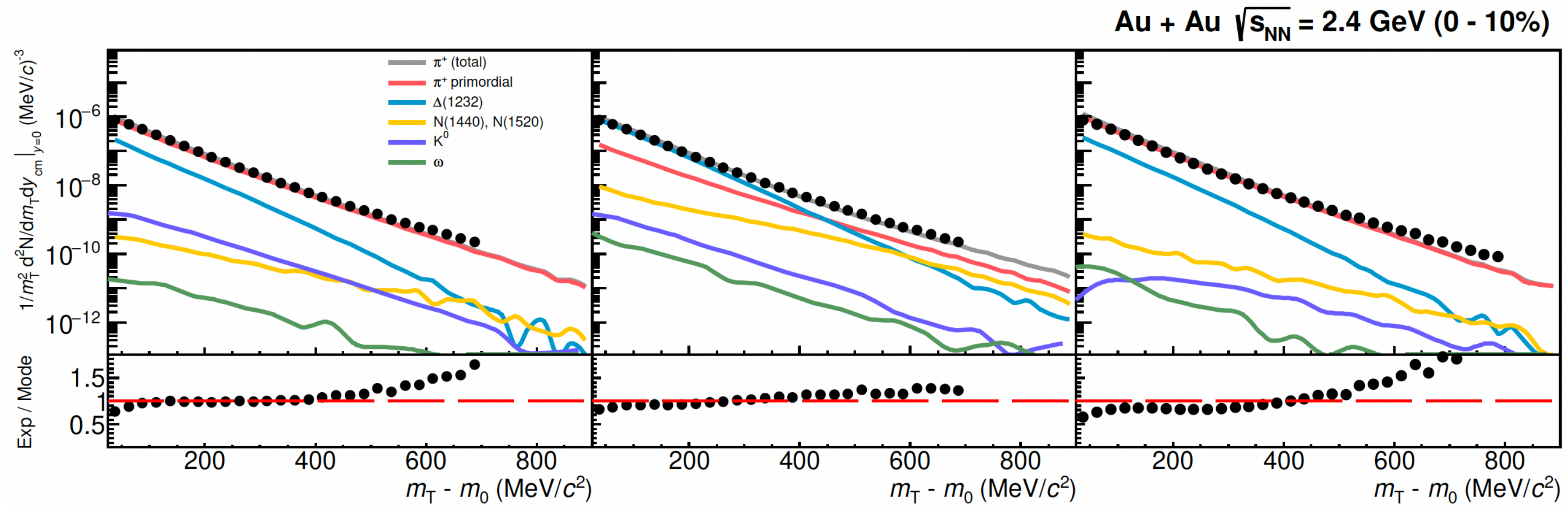
$S = 0$  and  $Q/B = 0.4$

We take for comparison transverse mass distributions of protons, + and - pions in five center-of-mass rapidity intervals:  
 [0:45;0:35],  
 [0:25;0:15],  
 [0:05; 0:05],  
 [0:15; 0; 25],  
 [0:35; 0:45]

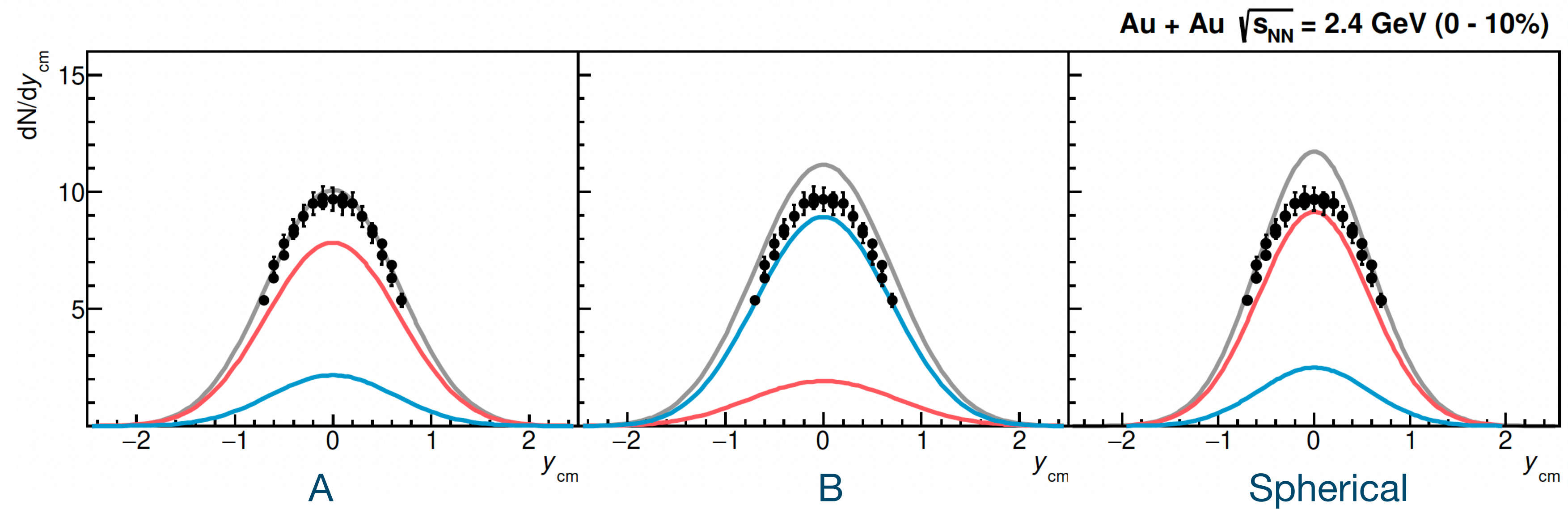
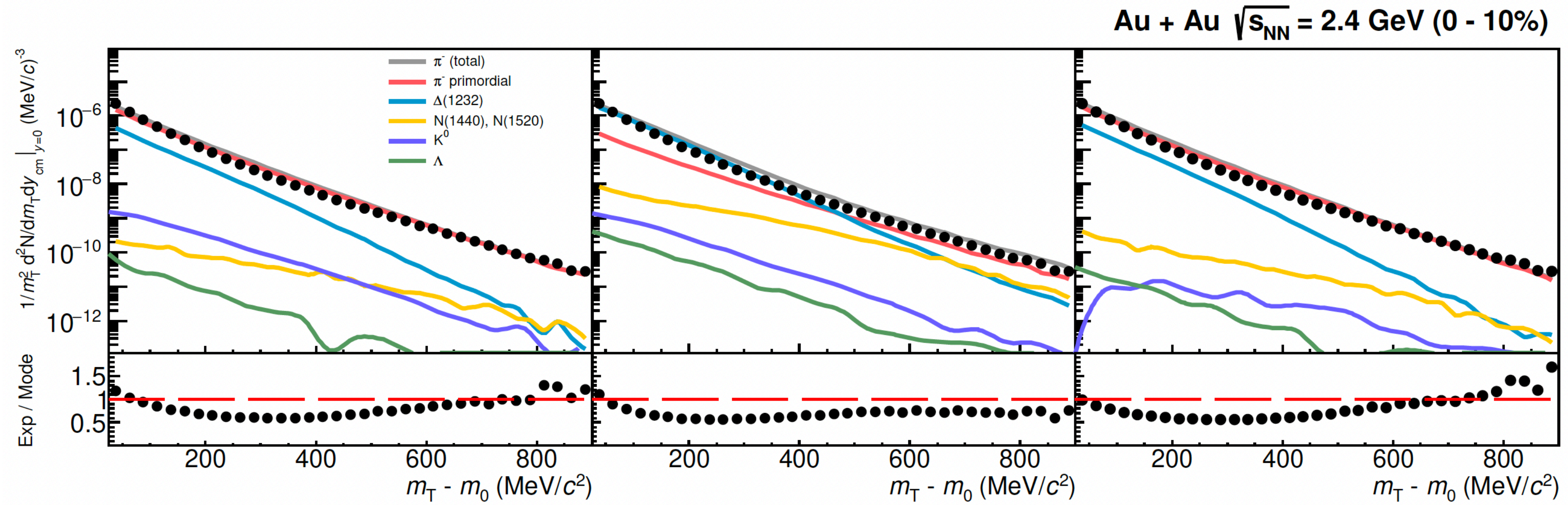
# RESULTS



# RESULTS



# RESULTS



# CONCLUSIONS

- We have **studied the rapidity and transverse mass spectra of protons and pions** produced in Au-Au collisions at 2.4 GeV and measured by HADES.
- We have found that they can be **well reproduced in a extended SR model** that assumes **single freeze-out of hadrons from a hypersurface spheroidal along beam direction**.
- Our framework modifies and extends RS approach by incorporation of the Hubble-like expansion of matter, inclusion of the resonance decays, and spheroidal deformation of the source.
- We have found that the presence of the **Delta resonance affects the spectra of pions**, while the contributions from other resonances can be neglected.
- The **obtained thermodynamic parameters agree well with the universal freeze-out curve established by other groups**.
- Our results bring **evidence for substantial thermalization of the matter produced in the few-GeV energy range and its nearly spherical expansion**.

**THANK YOU FOR YOUR ATTENTION.**



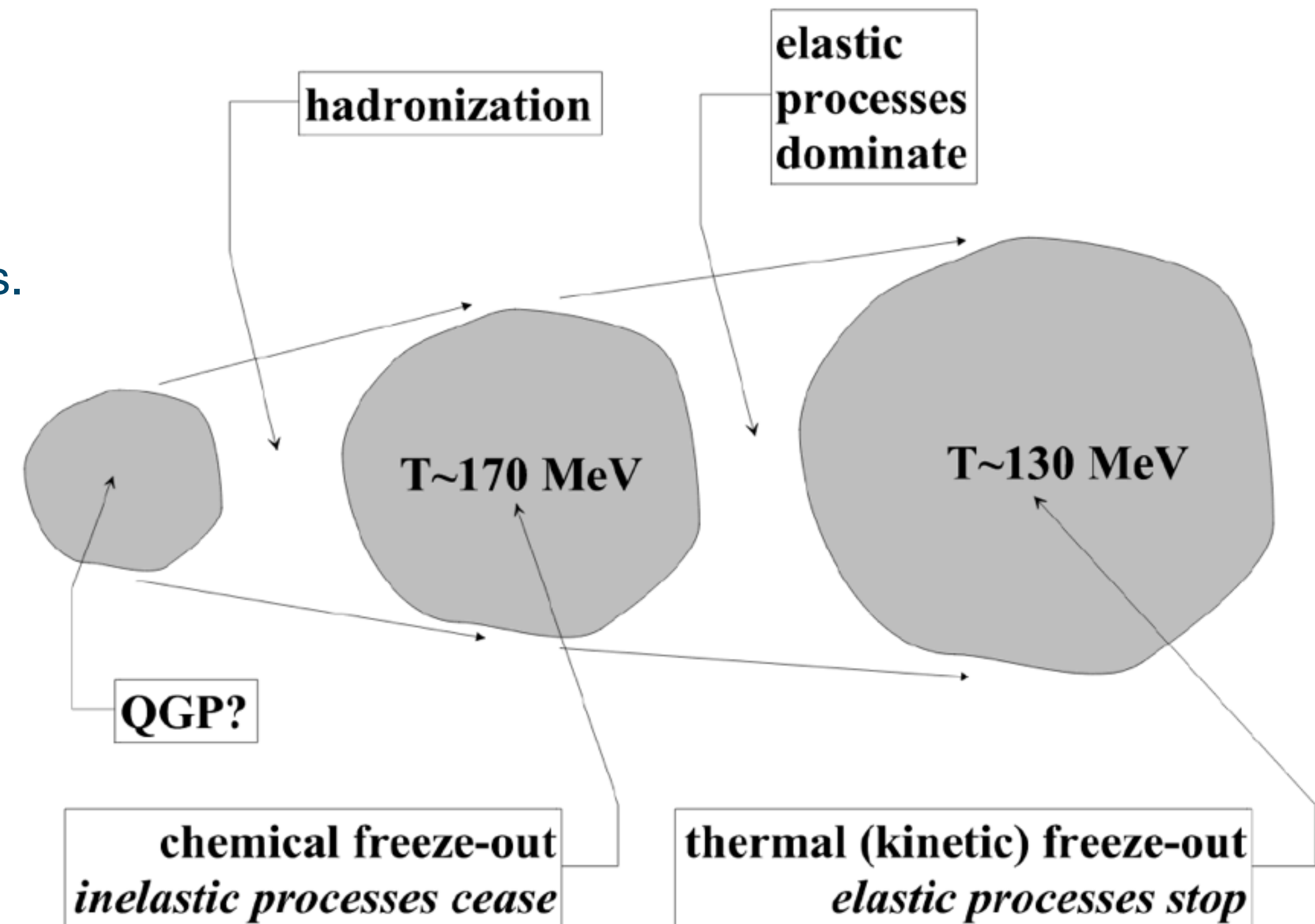
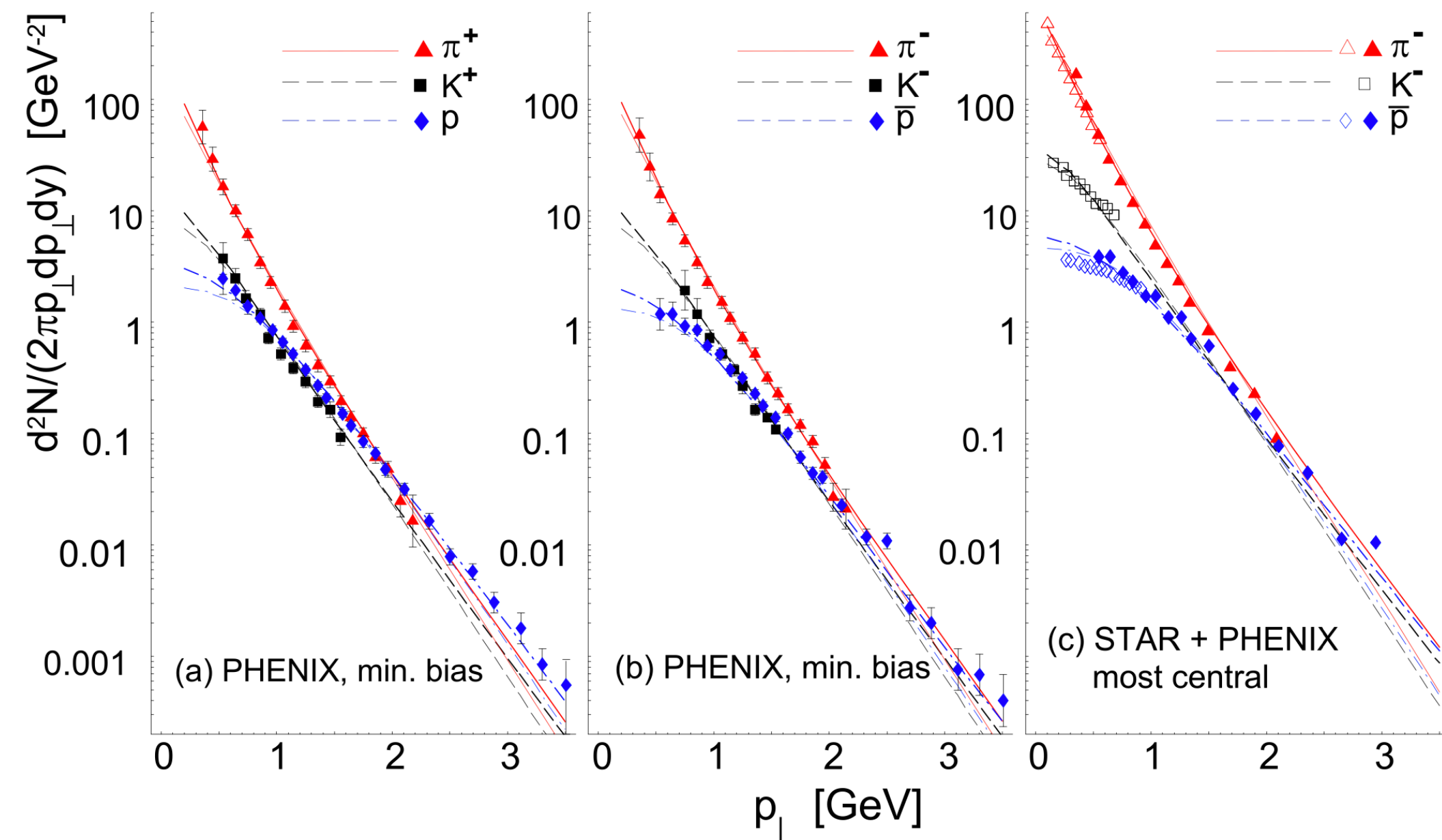
# SINGLE-FREEZE-OUT SCENARIO

1) Studies of ratios of hadron yields define **chemical freeze-out**.

2) Studies of spectra of hadrons define **kinetic freeze-out**.

We adopt the **single-freeze-out scenario** where chemical and kinetic freeze-outs coincide - successful at high energies.

*W. Broniowski, W. Florkowski, PRL 87, 272302 (2001)*



# DELTA RESONANCE TREATMENT

